

Optical signatures of Quantum Vacuum Nonlinearity in strong electromagnetic fields

(^{here} \dashrightarrow Heisenberg Euler eff. action
& photon prop. effects \rightarrow vacuum birefringence)

Helmholtz International Summer School

"QFT at the limits: from Strong Fields to Heavy Quarks"

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Felix Karbstein

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felix.karbstein@uni-jena.de

Introduction

Classical understanding: vacuum is empty

\equiv absence of any thing

(true / pure vacuum : no fields, boundary conditions, temperature)
 $\rightarrow \vec{E}, \vec{B}, \vec{g}, T \rightarrow 0, V \rightarrow \infty \quad \leftrightarrow \text{Lab}$

Vacuum + electromagnetic fields

$$\mathcal{L}_{MW} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

$$\mathcal{L}_j = -j_\mu A^\mu$$

$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$$

$$\mathcal{L}_{ED} = \mathcal{L}_{MW} + \mathcal{L}_j$$

$$S_{ED} = \int d^4x \mathcal{L}_{ED}$$

$$\mathcal{L}_{ED} = \mathcal{L}(A_\nu, \partial_\mu A_\nu) \quad \text{or equivalently} \quad \mathcal{L}(A_\nu, F_{\mu\nu})$$

\downarrow EOM

$$\frac{\delta S_{ED}}{\delta A_\nu} = 0 \quad \rightarrow \quad \frac{\partial \mathcal{L}}{\partial A_\nu} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} = 0$$

$$\leftrightarrow \frac{\partial \mathcal{L}}{\partial A_\nu} - \partial_\mu \frac{\partial \mathcal{L}}{\partial F_{\mu\nu}} = 0$$

$\left. \begin{array}{l} \downarrow \\ \leftarrow \end{array} \right\} \text{EOM}$

dh. für $\mathcal{L}_{MW} \rightarrow \partial_\mu F^{\mu\nu} = 0$ Maxwell Eqs.

\rightarrow superposition principle holds ;

if A_i^ν with associated $F_i^{\mu\nu}$ are solutions then clearly also $\sum_i A_i^\nu$.

Different in quantum vacuum ← vacuum of QFT
 here QED

The QVac is not empty but rather permeated by fluctuations / "virtual processes" of the fields of the considered theory. In QED : electrons & photons
 positrons

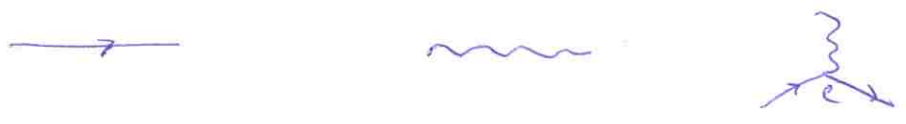
These virtual processes can be excited/influenced by external influences, e.g. strong (classical) electromagnetic fields or boundary conditions (→ Casimir type experiments)
Heisenberg & Euler '36
Casimir '48

→ response encodes the signal/imprint of the fluctuating fields.

Quantum electrodynamics: $(\hbar = c = 1)$

$$\mathcal{L}_{QED} = \bar{\Psi} (i\not{D} - m) \Psi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

$$\text{with } \not{D} = \gamma_{\mu} D^{\mu} \text{ \& } D^{\mu} = \partial^{\mu} - ie A^{\mu}$$



Ψ : 4-component complex Dirac spinor
 (anti-commuting Grassmann-valued field)

γ matrices satisfy $\{\gamma_{\mu}, \gamma_{\nu}\} = -2g_{\mu\nu}$

here $g_{\mu\nu} = \text{diag}(-1, +1, +1, +1)$

basic principle: "all processes that can be drawn can happen (if they are not forbidden)"

generically more suppressed with increasing loop order

Strong field QED \rightarrow towards the H.E. limit. ③
 In a next step we focus on QED in an external / classical field $A^\mu \rightarrow A_{cl}^\mu$ and (for the moment) completely ignore dynamical / quantized photons.

\rightarrow We are only interested in the dominant / leading order effect & it can be shown that virtual dynamical photons only contribute at higher loop order.

Hence, from now on we focus on

$$\mathcal{L} = \bar{\Psi} (i\not{D} - m) \Psi - \frac{1}{4} F^{\mu\nu} F_{\mu\nu}$$

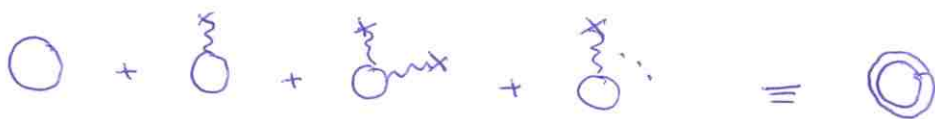
$$\text{where now } D^\mu = \partial^\mu - ie A_{cl}^\mu$$

$$F^{\mu\nu} = \partial^\mu A_{cl}^\nu - \partial^\nu A_{cl}^\mu$$

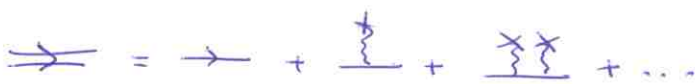


Here, we will only be interested in effective interactions among electromagnetic fields (\rightarrow lasers, "optical signatures" ^(as probe)) and do not consider situations / signatures with real e^+ / e^- (in the initial & final states) \rightarrow "~~optical~~"

Note that it is immediately clear that all possible Feynman diagrams (connected) which can be drawn for this theory are without external e^+ / e^- lines



with "dressed propagator"



In order to evaluate ③ explicitly we turn to the partition function \mathcal{Z} and (in a first step) integrate out the Dirac field providing us with an effective action without any expl. reference to the Dirac field.

To this end, recall that $\int \mathcal{D}\bar{\Psi} \int \mathcal{D}\Psi e^{\bar{\Psi} M \Psi} = \det M$.

$$\mathcal{Z}[A_{\mu}^M] \sim \int \mathcal{D}\bar{\Psi} \int \mathcal{D}\Psi e^{iS[A_{\mu}^M]} \quad (4)$$

$$= \int \mathcal{D}\bar{\Psi} \int \mathcal{D}\Psi e^{i \int d^4x \bar{\Psi} (i\not{D} - m) \Psi + iS_{MW}}$$

$$\left. \begin{aligned} & S_{MW} = \int d^4x \left(-\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right) \\ & = e^{iS_{MW}} \det(-\not{D} - im) = e^{i[S_{MW} - i \ln \det(-\not{D} - im)]} \\ & \sim e^{iS_{\text{eff}}[A_{\mu}^M]} \leftarrow \text{defined up to a constant} \end{aligned} \right\}$$

It is convenient to demand that $\mathcal{Z}[A_{\mu}^M = 0] = 1 \leftrightarrow S_{\text{eff}}[A_{\mu}^M = 0] = 0$

$$\rightarrow \text{In turn } \mathcal{Z}[A_{\mu}^M] = \frac{e^{iS_{MW}} \det(-\not{D} - im)}{\det(-\not{D} - im)}$$

$$= e^{iS_{MW}} \frac{\det(-i\not{D} + m)}{\det(-i\not{D} + m)}$$

$$\text{and } S_{\text{eff}}[A_{\mu}^M] = S_{MW} - i \ln \det(-i\not{D} + m) + i \ln \det(-i\not{D} + m)$$

$$\equiv S_{MW}[A_{\mu}^M] + S^1[A_{\mu}^M]$$

$$\text{with } S^1[A_{\mu}^M] = -i \ln \det(-i\not{D} + m) + i \ln \det(-i\not{D} + m)$$

$$" \textcircled{0} - \textcircled{0} "$$

In even spacetime dimensions we can define matrix γ_5 which fulfills $\{\gamma^{\mu}, \gamma_5\} = 0$; $\gamma_5^2 = \mathbb{1}$

$$(\gamma_5 = i\gamma^0\gamma^1\gamma^2\gamma^3)$$

$$\text{Therewith, } \det(-i\not{D} + m) = \det[(-i\not{D} + m)\gamma_5^2]$$

$$= \det[\gamma_5 (+i\not{D} + m)\gamma_5] = \det(i\not{D} + m)$$

$$\Leftrightarrow \det^2(-i\not{D} + m) = \det(-i\not{D} + m) \det(i\not{D} + m) \\ = \det(\not{D}^2 + m^2)$$

$$\rightarrow \ln \det(-i\not{D} + m) = \frac{1}{2} \ln \det(\not{D}^2 + m^2)$$

Such that $S^1 = -\frac{i}{2} \ln \det(\not{D}^2 + m^2) + \frac{i}{2} \ln \det(\not{D}^2 + m^2)$

$$\not{D}^2 = D_\mu D_\nu \gamma^\mu \gamma^\nu = D_\mu D_\nu \frac{1}{2} (\{\gamma^\mu, \gamma^\nu\} + [\gamma^\mu, \gamma^\nu])$$

$$\left\{ \begin{aligned} \{\gamma^\mu, \gamma^\nu\} &= -2g^{\mu\nu} \mathbb{1} \\ [\gamma^\mu, \gamma^\nu] &\equiv -2i\sigma^{\mu\nu} \leftarrow \text{antisym.} \end{aligned} \right.$$

$$= -D^2 - i\sigma^{\mu\nu} D_\mu D_\nu \\ = -D^2 - \frac{i}{2} \sigma^{\mu\nu} \underbrace{[D_\mu, D_\nu]}_{= \frac{e}{i} F_{\mu\nu}}$$

$$= -D^2 - \frac{e}{2} \sigma^{\mu\nu} F_{\mu\nu}$$

and hence $S^1 = -\frac{i}{2} \ln \det(-D^2 + m^2 - \frac{e}{2} \sigma^{\mu\nu} F_{\mu\nu}) \\ + \frac{i}{2} \ln \det(-\partial^2 + m^2)$

$$\int \ln \det \dots = \text{Tr} \ln \dots \quad (\ln \prod_i \lambda_i = \sum_i \ln \lambda_i)$$

$$\rightarrow S^1 = -\frac{i}{2} \text{Tr} \ln(-D^2 + m^2 - \frac{e}{2} \sigma^{\mu\nu} F_{\mu\nu}) + \frac{i}{2} \text{Tr} \ln(-\partial^2 + m^2)$$

In the next step we employ the "proper time representation" of the log; Schwinger '51

$$\ln M - \ln M_0 = \lim_{\Lambda \rightarrow \infty} \left\{ - \int_{1/\Lambda^2}^{\infty} \frac{dT}{T} (e^{-MT} - e^{-M_0 T}) \right\}$$

for $\text{Re}\{M, M_0\} > 0$; sufficient here, cf. below.

Here, $\Lambda :=$ regulator for UV-divergences \leftrightarrow large energies

⌈ This follows straightforwardly from

$$\lim_{\Lambda \rightarrow \infty} \int_{1/\Lambda^2}^{\infty} dT e^{-\tilde{M}T} = \frac{1}{\tilde{M}} \text{ by integration } \int_{M_0}^M d\tilde{M}$$

and exchange of int.]

$$\rightarrow S^1 = \frac{i}{2} \int_{1/\Lambda^2}^{\infty} \frac{dT}{T} \left\{ \text{Tr} e^{-(-D^2+m^2 - \frac{e}{2} G^{\mu\nu} F_{\mu\nu})T} - \text{Tr} e^{-(-\partial^2+m^2)T} \right\}$$

where $\lim_{\Lambda \rightarrow \infty}$ is implicitly understood (to be taken at the very end of the calculation).

----- Ende 1. Vorl. -----

This expression is still valid for arbitrary A_{cl}^M fields (At least $\text{Tr}\{\dots\}$) can be evaluated explicitly for several special cases:

- constant electromagnetic fields, arbitrary orientation
- plane wave "null" fields $A_{cl}^M \sim \cos(x^M x_\mu)$ with $x^2=0$ and $\vec{E} \perp \vec{B}$, $|\vec{E}|=|\vec{B}|$. $\rightarrow S^1=0$
- 1d magnetic field inhomogeneity $\sim \text{sech}^2(x) = \frac{1}{\text{ch}^2(x)}$
- starting point of worldline formalism to QED
- numerics

Subsequently we stick to a constant magnetic field

→ explicit evaluation of $S^1[B]$

choose $A_{cl}^M = (0, 0, Bx, 0) \rightarrow \vec{B} = B \vec{e}_z$

→ $F_{12} = -F_{21} = B$; all other components vanish

⌈

$$F^{\mu\nu} = \begin{pmatrix} 0 & \vec{E}^T \\ \vec{E} & \begin{pmatrix} 0 & B_3 & -B_2 \\ -B_3 & 0 & B_1 \\ B_2 & -B_1 & 0 \end{pmatrix} \end{pmatrix}$$

⌋

To this end we use the following conventions for the γ -Matrices

(7)

$$\gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}, \quad \gamma^i = \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix}, \quad G^{\mu\nu} = \frac{i}{2} [\gamma^\mu, \gamma^\nu]$$

with Pauli matrices σ^i ; $\sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$, $\sigma^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$, $\sigma^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$

$$\frac{1}{2} [\sigma^i, \sigma^j] = i \epsilon_{ijk} \sigma^k$$

$$\rightarrow G^{12} = \frac{i}{2} \begin{pmatrix} [\sigma^2, \sigma^1] & 0 \\ 0 & [\sigma^2, \sigma^1] \end{pmatrix} = \begin{pmatrix} \sigma^3 & 0 \\ 0 & \sigma^3 \end{pmatrix} = \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & 1 & \\ & & & -1 \end{pmatrix}$$

$$\rightarrow \frac{e}{2} F^{\mu\nu} G_{\mu\nu} = e F^{12} G_{12} = e B \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & 1 & \\ & & & -1 \end{pmatrix}$$

$$D^2 = (\partial^\mu - ie A_{\alpha}^\mu) (\partial_\mu - ie A_{\alpha, \mu})$$

$$= -\partial_+^2 + \partial_-^2 + D_\perp^2 \quad \text{with } D_\perp^2 = \sum_{j=1}^2 (\partial^j - ie A_{\alpha}^j) (\partial_j - ie A_{\alpha, j})$$

$$\rightarrow D_\perp^2 = \partial^j \partial_j - ie (\partial_j A_{\alpha}^j) - 2ie A_{\alpha}^j \partial_j - e^2 A_{\alpha}^j A_{\alpha, j}$$

$$= \partial_x^2 + \partial_y^2 - 2ie B_x \partial_y - (e B_x)^2$$

$$= \partial_x^2 + (\partial_y - ie B_x)^2$$

In order to perform the trace, we need the eigenvalues of D^2 .

\rightarrow eigenvalues of $-(-\partial_+^2 + \partial_-^2)$ are plane wave eigenvalues

$$-p_+^2 + p_-^2 \quad \& \text{ after rotation to Euclidean } p_+ \rightarrow ip_\tau$$

$$+ p_\tau^2 + p_z^2$$

\rightarrow eigenvalues for $-D_\perp^2$?

$$-D_\perp^2 f(x, y) = \lambda f(x, y)$$

$$-[\partial_x^2 + (\partial_y - ie B_x)^2] f(x, y) = \lambda f(x, y)$$

$$\text{ansatz } f(x, y) = e^{ip_y y} f(x)$$

$$- [\partial_x^2 - (p_y - eBx)^2] f(x) = \lambda f(x)$$

$$- [\partial_x^2 - (eB(x - \frac{p_y}{eB}))^2] f(x) = \lambda f(x)$$

defining $g = \sqrt{eB} (x - \frac{p_y}{eB})$, $\frac{\partial}{\partial x} = \frac{\partial g}{\partial x} \frac{\partial}{\partial g} = \sqrt{eB} \frac{\partial}{\partial g}$

$$-eB [\partial_g^2 - g^2] \tilde{f}(g) = \lambda \tilde{f}(g)$$

$$[(-i\partial_g)^2 + g^2] \tilde{f}(g) = \frac{\lambda}{eB} \tilde{f}(g)$$

c.f. harmonic oscillator in quantum mechanics (1d)

$$\hat{H} = \frac{\hat{p}^2}{2m} + \frac{1}{2} m\omega^2 \hat{x}^2 , \hat{p} = -i\hbar \partial_x$$

→ spectrum $\hat{H} |\Psi\rangle = E |\Psi\rangle$

(pos. space $\hat{H} \Psi(x) = E \Psi(x)$)

→ eigenvalues $E_n = \hbar\omega (n + 1/2)$, $n \in \mathbb{N}_0^+$

→ $2\hat{H} \Psi(x) = [\hat{p}^2 + \hat{x}^2] \Psi(x) = 2E \Psi(x)$

for $\hbar = m = \omega = 1$

↔ $\frac{\lambda_n}{eB} = 2E_n = 2(n + \frac{1}{2})$

→ $\lambda_n = 2eB (n + \frac{1}{2})$ Landau levels

Hence, $\text{Tr} \rightarrow i \int \frac{dp_x}{2\pi} \int \frac{dp_z}{2\pi} \sum_{n=0}^{\infty} g(n) \text{tr}_x$ } 1st term of S^1 .

↑
density of states for Landau levels

$-D^2 + m^2 \rightarrow p_x^2 + p_z^2 + 2eB(n + \frac{1}{2}) + m^2$

(9)

$$\left. \begin{aligned} \text{while Tr} \rightarrow 4i \int \frac{dp_z}{2\pi L} \int \frac{dp_x}{2\pi L} \int \frac{dp_y}{2\pi L} \int \frac{dp_z}{2\pi L} \\ -\partial^2 + m^2 \rightarrow p_\tau^2 + p_x^2 + p_y^2 + p_z^2 + m^2 \end{aligned} \right\} \begin{array}{l} \text{2nd} \\ \text{term of } S^1 \end{array}$$

$g(n)$ can be worked out from limit $eB \rightarrow 0$

(and accounting for $S_{\text{eff}} \approx$ extensive quantity $\sim V$)

$$\begin{aligned} \lim_{eB \rightarrow 0} \sum_{n=0}^{\infty} g(n) &\stackrel{!}{=} \int \frac{dp_x}{2\pi L} \int \frac{dp_y}{2\pi L} = \frac{L^2}{(2\pi)^2} \int_0^{\infty} dp_{\perp} p_{\perp} \int_0^{2\pi} d\varphi \\ &= \frac{L^2}{4\pi} \int_0^{\infty} dp_{\perp}^2 = \frac{L^2}{4\pi} \lim_{eB \rightarrow 0} \sum_{n=0}^{\infty} \Delta p_{\perp}^2 \end{aligned}$$

$$\text{as } p_{\perp}^2 = p_x^2 + p_y^2 \Leftrightarrow \Delta p_{\perp}^2 = 2eB$$

$$= \frac{L^2}{2\pi} \lim_{eB \rightarrow 0} \sum_{n=0}^{\infty} eB$$

we obtain $g(n) = \frac{L^2}{2\pi} eB$.

$$\text{And with } \int \frac{dp}{2\pi L} e^{-p^2 T} = L \frac{1}{\sqrt{4\pi T}}$$

$$\begin{aligned} \sum_{n=0}^{\infty} e^{-eBT(2n+1)} &= \frac{e^{-eBT}}{1 - e^{-2eBT}} = \frac{1}{e^{eBT} - e^{-eBT}} \\ &= \frac{1}{2 \sinh(eBT)} \end{aligned}$$

$$\text{tr}_g e^{\frac{e}{2} G^{\mu\nu} F_{\mu\nu} T} \underset{\substack{\uparrow \\ \text{B field}}}{=} \text{tr}_g e^{eBT \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & 1 & \\ & & & -1 \end{pmatrix}}$$

$$= 2 (e^{eBT} + e^{-eBT}) = 4 \cosh(eBT)$$

we finally obtain

$$S^1 = -\frac{1}{2} \int_{-1/2}^{\infty} \frac{dT}{T^2} e^{-m^2 T} \frac{L^2}{4\pi} \left\{ \frac{L^2}{2\pi} eB 2 \cosh(eBT) - \frac{L^2}{\pi} \frac{1}{T} \right\}$$

$$S^1 = - \frac{L^4}{8\pi^2} \int_{1/\Lambda^2}^{\infty} \frac{dT}{T^3} e^{-m^2 T} \left\{ (eBT) \coth(eBT) - 1 \right\}$$

\cong Heisenberg-Euler action (unren.) for purely B-field

Note that $x \coth x - 1 = \frac{1}{3} x^2 - \frac{1}{45} x^4 + \mathcal{O}(x^6)$

\leadsto log-type divergence for $\Lambda \rightarrow \infty$

We write (subtracting & adding)

$$S^1 = - \frac{L^4}{8\pi^2} \int_{1/\Lambda^2}^{\infty} \frac{dT}{T^3} e^{-m^2 T} \left\{ eBT \coth(eBT) - \frac{1}{3} (eBT)^2 - 1 \right\}$$

$$- \frac{L^4}{24\pi^2} (eB)^2 \int_{1/\Lambda^2}^{\infty} \frac{dT}{T} e^{-m^2 T}$$

subst. $m^2 T = \tilde{T} \rightarrow \int_{(m/\Lambda)^2}^{\infty} \frac{d\tilde{T}}{\tilde{T}} e^{-m^2 \tilde{T}}$

$v \ll 1$
fixed
 \downarrow
 $\int_v^{\infty} \frac{d\tilde{T}}{\tilde{T}} e^{-m^2 \tilde{T}} + \int_{(m/\Lambda)^2}^v \frac{d\tilde{T}}{\tilde{T}} (1 - \tilde{T} + \mathcal{O}(\tilde{T}^2))$

$$= \ln \frac{\Lambda^2}{m^2} + \text{const.} + \mathcal{O}\left(\frac{m^2}{\Lambda^2}\right)$$

\rightarrow the divergent contribution is $\sim B^2$

cf. $S_{MW} = -L^4 \frac{1}{4} F_{\mu\nu} F^{\mu\nu} = -L^4 \frac{1}{2} (B^2 - E^2) \underset{E=0}{=} -L^4 \frac{1}{2} B^2 \sim B^2$

$\rightarrow S_{\text{eff}} = S_{MW} + S^1$; $\mathcal{L}_{\text{eff}} = \frac{S_{\text{eff}}}{L^4}$

$$\mathcal{L}_{\text{eff}}(B) = - \frac{1}{2} B^2 \left\{ 1 + \frac{e^2}{12\pi^2} \left[\ln \frac{\Lambda^2}{m^2} + \text{const.} \right] \right\}$$

$$- \frac{e^2}{24\pi^2} B^2 \ln \frac{m^2}{m^2}$$

$$- \frac{1}{8\pi^2} \int_0^{\infty} \frac{dT}{T^3} e^{-m^2 T} \left\{ eBT \coth(eBT) - \frac{1}{3} (eBT)^2 - 1 \right\}$$

Recall that so far we have worked with unrenormalized fields and not bothered about renormalization.

→ Renormalization: \rightarrow the term $\sim B^2$ should match the physical (measurable) Maxwell term (11)

To this end we introduce wavefunction renormalization

$$Z^{-1} \equiv 1 + \frac{e^2}{12\pi^2} \left[\ln \frac{\Lambda^2}{\mu^2} + \text{const.} \right]$$

and define $B_R^2 = Z^{-1} B^2$. (Note that $A_{cl}^M \sim B \leftrightarrow A_{cl,R}^M = Z^{-1/2} A_{cl}^M$.)

With this rescaling, the vertex/interaction in the Lagrangian of the microscopic theory $e A_{cl}^M \bar{\Psi} \gamma_\mu \Psi \rightarrow e Z^{1/2} A_{cl,R}^M \bar{\Psi} \gamma_\mu \Psi$.

Demanding it to be given by $e_R A_{cl,R}^M \bar{\Psi} \gamma_\mu \Psi$ we infer $e_R^2 = Z e^2$.

In turn, we have

$$B_R = Z^{-1/2} B = B_R(\mu)$$

$$e_R = Z^{1/2} e = e_R(\mu)$$

but $e_R B_R = e B$, μ -indep.

Correspondingly, the renormalized H.E. Lagrangian reads

$$\mathcal{L}_{\text{eff}}^{\text{ren}}(B) = -\frac{1}{2} B_R^2 - \frac{e_R^2}{24\pi^2} B_R^2 \ln \frac{\Lambda^2}{\mu^2}$$

$$- \frac{1}{8\pi^2} \int_0^\infty \frac{dT}{T^3} e^{-m^2 T} \left\{ e_R B_R T \coth(e_R B_R T) - \frac{1}{3} (e_R B_R T)^2 - 1 \right\}.$$

Conventionally the "on-shell" renormalization condition $\mu \equiv m$ is adopted

$$\frac{e_R^2(\mu=m)}{4\pi} = \alpha_R(\mu=m) \approx \frac{1}{137}$$

Ende 2. Vorl.

$$\xrightarrow{\mu=m} \mathcal{L}_{\text{eff}}^{\text{ren}}(B) = -\frac{1}{2} B_R^2 - \frac{1}{8\pi^2} \int_0^\infty \frac{dT}{T^3} e^{-m^2 T} \left\{ e_R B_R T \coth(e_R B_R T) - \frac{1}{3} (e_R B_R T)^2 - 1 \right\}$$

It can be shown that for arbitrary/generic constant electromagnetic

fields $\mathcal{L}_{\text{eff}} \equiv \mathcal{L}_{\text{eff}}(\mathcal{F}, \mathcal{G}^2)$ with $\mathcal{F} = \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$, $\mathcal{G} = \frac{1}{4} F_{\mu\nu} \tilde{F}^{\mu\nu}$

$$\text{with } \tilde{F}^{\mu\nu} = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta}$$

$$\left(\mathcal{F} = \frac{1}{2} (\vec{B}^2 - \vec{E}^2), \quad \mathcal{G} = -\vec{E} \cdot \vec{B} \right)$$

In order to write it compactly we moreover define

$$a = \left(\sqrt{\mathcal{F}^2 + \mathcal{G}^2} - \mathcal{F} \right)^{1/2}, \quad b = \left(\sqrt{\mathcal{F}^2 + \mathcal{G}^2} + \mathcal{F} \right)^{1/2}.$$

The result is

$$\mathcal{L}_{\text{eff}}^{\text{ren}}(\mathcal{F}, \ell_f^2) = -\mathcal{F} - \frac{1}{8\pi^2} \int_0^\infty \frac{dT}{T^3} e^{-m^2 T} \left\{ (eT)^2 ab \coth(Teb) \cot(eTa) + \frac{1}{3} (eT)^2 (a^2 - b^2) - 1 \right\}$$

→ contains all orders in \mathcal{F}, ℓ_f^2

→ is even in $e \leftrightarrow$ charge conjugation invariance of QED Furry '37

To lowest order - counting $\sigma(\mathcal{F}) = \sigma(\ell_f) = \sigma(F^2)$ - we obtain

$$\mathcal{L}_{\text{eff}} = -\mathcal{F} + \underbrace{\frac{8}{45} \frac{\alpha^2}{m^4} \mathcal{F}^2}_{\equiv c_1} + \underbrace{\frac{14}{45} \frac{\alpha^2}{m^4} \ell_f^2}_{\equiv c_2} + \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^6\right)$$

Application(s)

Let us now study light propagation

The results derived for constant electromagnetic fields can be adopted for "weakly varying" fields also.

(At least for weak fields) The QED scale is $m \leftrightarrow$ length scale

$$\lambda_c = \frac{1}{m} = 3.8 \cdot 10^{-13} \text{ m. Consider derivative expansion around}$$

const. - field result → derivatives are rendered dimensionless by

$$\lambda_c \sim (\lambda_c \partial_x)^n \sim \left(\frac{\omega}{m}\right)^n \text{ where } \omega \text{ is the typical frequency scale of variation of the considered field configuration. For}$$

$\left(\frac{\omega}{m}\right) \ll 1$ these can be neglected and we can employ the

substitution $F^{\mu\nu} \rightarrow F^{\mu\nu}(x)$ in constant-field-result for \mathcal{L}_{eff} .

Let us stick to this assumption and consider the equations of motion (cf. begin of lecture)

$$\text{Here } \mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{eff}}(F^{\mu\nu}) = \mathcal{L}_{\text{eff}}(\mathcal{F}, \ell_f^2) \quad ; \quad \mathcal{L}_{\text{eff}} = -\mathcal{F} + \mathcal{L}^1$$

$$\begin{aligned} \rightarrow \partial_\mu \frac{\partial \mathcal{L}_{\text{eff}}}{\partial F_{\mu\nu}} &= 0 \quad \rightarrow \partial_\mu \left(F^{\mu\nu} - \frac{\partial \mathcal{L}^1}{\partial \mathcal{F}} F^{\mu\nu} - \frac{\partial \mathcal{L}^1}{\partial \ell_f} \tilde{F}^{\mu\nu} \right) \\ &= 0 \end{aligned}$$

$$\frac{\partial \mathcal{L}^1}{\partial \mathcal{F}} = 2c_1 \mathcal{F} \left(1 + \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^2\right) \right) \quad ; \quad \frac{\partial \mathcal{L}^1}{\partial \ell_f} = 2c_2 \ell_f \left(1 + \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^2\right) \right)$$

$$\rightarrow \partial_\mu \left\{ (1 - 2c_1 \mathcal{F}) F^{\mu\nu} - 2c_2 \mathcal{G} \tilde{F}^{\mu\nu} \right\} = \mathcal{O}(F^5)$$

In next step we decompose $F^{\mu\nu}(x) \rightarrow F^{\mu\nu} + f^{\mu\nu}(x)$
 \uparrow const. \uparrow probe

and linearize in $f \ll F$.

Note that $\mathcal{F} \rightarrow \mathcal{F} + \frac{1}{2} F_{\mu\nu} f^{\mu\nu} + \mathcal{O}(f^2)$
 $\mathcal{G} \rightarrow \mathcal{G} + \frac{1}{2} \tilde{F}_{\mu\nu} f^{\mu\nu} + \mathcal{O}(f^2)$

Generic $F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$ fulfills Bianchi Id.

$$\partial^\alpha F^{\beta\gamma} + \partial^\beta F^{\gamma\alpha} + \partial^\gamma F^{\alpha\beta} = 0 \leftarrow \text{checked by insertion}$$

$$\Leftrightarrow \partial_\mu \tilde{F}^{\mu\nu} = 0$$

Bianchi
 $\rightarrow \partial_\mu \left\{ (1 - 2c_1 \mathcal{F}) F^{\mu\nu} \right\} - 2c_2 (\partial_\mu \mathcal{G}) \tilde{F}^{\mu\nu} = \mathcal{O}(F^5)$

lin.
 $\rightarrow (1 - 2c_1 \mathcal{F}) \partial_\mu f^{\mu\nu} - 2c_2 \underbrace{(\partial_\mu \mathcal{G})}_{\rightarrow 0} \tilde{f}^{\mu\nu}$
 $- 2c_1 \partial_\mu \left(\frac{1}{2} F_{\alpha\beta} f^{\alpha\beta} \right) F^{\mu\nu} - 2c_2 \partial_\mu \left(\frac{1}{2} \tilde{F}_{\alpha\beta} f^{\alpha\beta} \right) \tilde{F}^{\mu\nu} = 0$

$$\Leftrightarrow (1 - 2c_1 \mathcal{F}) \partial_\mu f^{\mu\nu} - c_1 F_{\alpha\beta} F^{\mu\nu} \partial_\mu f^{\alpha\beta} - c_2 \tilde{F}_{\alpha\beta} \tilde{F}^{\mu\nu} \partial_\mu f^{\alpha\beta} = 0$$

lin. equations of motion (pos. space; $f \equiv f(x)$)

In order to solve them we first go to momentum space

$$a^\mu(x) = \int \frac{d^4k}{(2\pi)^4} e^{ikx} a^\mu(k)$$

$$\rightarrow f^{\mu\nu}(x) = i \int \frac{d^4k}{(2\pi)^4} e^{ikx} [k^\mu a^\nu(k) - k^\nu a^\mu(k)]$$

and work in Lorenz gauge $\partial_\mu a^\mu(x) = 0 \Leftrightarrow k_\mu a^\mu(k) = 0$

~~$\rightarrow k_\mu f^{\mu\nu}(x) = i$~~

$$\rightarrow \partial_\mu f^{\mu\nu}(x) = - \int \frac{d^4k}{(2\pi)^4} e^{ikx} [k^2 a^\nu(k)]$$

$$(1-2c_1 \mathcal{F}) k^2 a^\nu(k) - c_1 F_{\alpha\beta} F^{\mu\nu} k_\mu [k^\alpha a^\beta(k) - k^\beta a^\alpha(k)] - c_2 \tilde{F}_{\alpha\beta} \tilde{F}^{\mu\nu} k_\mu [k^\alpha a^\beta(k) - k^\beta a^\alpha(k)] = 0$$

$$F_{\alpha\beta} = -F_{\beta\alpha}$$

$$\rightarrow \boxed{(1-2c_1 \mathcal{F}) k^2 a^\nu(k) - 2c_1 F_{\alpha\beta} F^{\mu\nu} k_\mu k^\alpha a^\beta(k) - 2c_2 \tilde{F}_{\alpha\beta} \tilde{F}^{\mu\nu} k_\mu k^\alpha a^\beta(k) = 0} \quad (*)$$

≅ Eq. of Mot. of type $M^\nu{}_\mu a^\mu = 0$

→ nontrivial solutions?

Strategy: use ansatz $a_1^\mu \sim F^{\mu\nu} k_\nu \equiv (Fk)^\mu = (\vec{k} \cdot \vec{E}, \vec{k} \times \vec{B} + \omega \vec{E})$
 $a_2^\mu \sim \tilde{F}^{\mu\nu} k_\nu \equiv (\tilde{F}k)^\mu = (\vec{k} \cdot \vec{B}, -\vec{k} \times \vec{E} + \omega \vec{B})$

and use contraction ids.

$$F^{\mu\alpha} F^\nu{}_\alpha - \tilde{F}^{\mu\alpha} \tilde{F}^\nu{}_\alpha = 2\mathcal{F} g^{\mu\nu}$$

$$F^{\mu\alpha} \tilde{F}^\nu{}_\alpha = \tilde{F}^{\mu\alpha} F^\nu{}_\alpha = \mathcal{G} g^{\mu\nu}$$

obviously $k_\mu a_i^\mu = 0$ due to symmetry

$$a_1^\mu a_{2\mu} = F^{\mu\nu} k_\nu \tilde{F}_\mu{}^\alpha k_\alpha = \mathcal{G} g^{\nu\alpha} k_\nu k_\alpha = \mathcal{G} k^2$$

From (*) we infer $k^2 = 0 + \mathcal{O}((\frac{eF}{m^2})^2) \propto$

With Maxwell's eqs. in vac, the electric/magnetic fields associated with a_i^μ read

$$\left. \begin{aligned} \vec{e}_i &= -\vec{\nabla} a_i^0 - \partial_+ \vec{a}_i \\ \vec{b}_i &= \vec{\nabla} \times \vec{a}_i \end{aligned} \right\} \text{in position space}$$

$$\rightarrow \left. \begin{aligned} \vec{e}_i(k) &\sim -\vec{k} a_i^0(k) + \omega \vec{a}_i(k) \\ \vec{b}_i(k) &\sim \vec{k} \times \vec{a}_i(k) \end{aligned} \right\} \text{in mom. space } k^0 = \omega$$

$$\vec{e}_1(k) \sim -\vec{k} (\vec{k} \cdot \vec{E}) + \omega (\vec{k} \times \vec{B}) + \omega^2 \vec{E}$$

$$\vec{b}_1(k) \sim \omega \vec{k} \times \vec{E} - \vec{k}^2 \vec{B} + \vec{k} (\vec{k} \cdot \vec{B})$$

$$\vec{e}_2(k) \sim -\vec{k} (\vec{k} \cdot \vec{B}) - \omega (\vec{k} \times \vec{E}) + \omega^2 \vec{B}$$

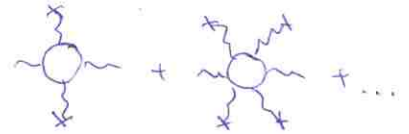
$$\vec{b}_2(k) \sim \omega \vec{k} \times \vec{B} + \vec{k}^2 \vec{E} - \vec{k} (\vec{k} \cdot \vec{E})$$

$$(*) \quad \boxed{(1-2c_1 \mathcal{F}) k^2 a^\nu(k) + 2c_1 (Fk)^\nu k^\alpha F_{\alpha\beta} a^\beta(k) + 2c_2 (\tilde{F}k)^\nu k^\alpha \tilde{F}_{\alpha\beta} a^\beta(k) = 0} \quad (\square)$$

$$\underline{a_1}: (1-2c_1 \mathcal{F}) k^2 (Fk)^\nu - 2c_1 (Fk)^\nu (Fk)^2 + 2c_2 (\tilde{F}k)^\nu k^\alpha k^\beta (-1) \eta_{\alpha\beta} g_{\alpha\beta} = 0$$

$$\Leftrightarrow [(1-2c_1 \mathcal{F}) k^2 - 2c_1 (Fk)^2] (Fk)^\nu - 2c_2 \eta_{\alpha\beta} k^2 (\tilde{F}k)^\nu = 0$$

solved with ansatz $k^2 = 0 + \delta_1$
 $= \alpha \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^2\right)$



\rightarrow can be systematically (recursively) extended to higher orders

$$\rightarrow \left[\underbrace{(1-2c_1 \mathcal{F}) \delta_1}_{\propto \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^4\right)} - 2c_1 (Fk)^2 \right] (Fk)^\nu - \alpha \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^4\right) (\tilde{F}k)^\nu = 0$$

$$\rightarrow \delta_1 = 2c_1 (Fk)^2 + \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^4\right) \propto$$

$$\underline{a_2}: (1-2c_1 \mathcal{F}) k^2 (\tilde{F}k)^\nu - 2c_1 (Fk)^\nu k^\alpha k^\beta \eta_{\alpha\beta} g_{\alpha\beta} - 2c_2 (\tilde{F}k)^\nu (\tilde{F}k)^2 = 0$$

$$\left. \begin{aligned} (\tilde{F}k)^2 &= \tilde{F}^{\mu\nu} k_\nu \tilde{F}_\mu^\alpha k_\alpha = k_\nu k_\alpha (F^{\mu\nu} F_\mu^\alpha - 2g^{\nu\alpha} \mathcal{F}) \\ &= (Fk)^2 - 2\mathcal{F} k^2 \end{aligned} \right\}$$

$$\Leftrightarrow [(1-2c_1 \mathcal{F} + 2c_2 \mathcal{F}) k^2 - 2c_2 (Fk)^2] (\tilde{F}k)^\nu - 2c_1 \eta_{\alpha\beta} k^2 (Fk)^\nu = 0$$

$$\rightarrow \delta_2 = 2c_2 (Fk)^2 + \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^4\right) \propto$$

the dispersion relations for modes $a_i^\mu(k)$ are given by

$$\boxed{k^2 = 2c_i (Fk)^2 + \alpha \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^4\right)}$$

$$(Fk)^2 = \omega^2 \vec{E}^2 + \vec{k}^2 \vec{B}^2 - (\vec{k} \cdot \vec{E})^2 - (\vec{k} \cdot \vec{B})^2 + 2(\vec{k} \times \vec{B}) \cdot \vec{E} \omega$$

$$\rightarrow \text{obviously } (Fk)^2 = (Fk)^2|_{k^2=0} + F^2 \mathcal{O}\left(\left(\frac{eF}{m^2}\right)^2\right)$$

$$\begin{aligned}
 (Fk)^2 |_{k^2=0} &= (\vec{k} \times \vec{E})^2 + (\vec{k} \times \vec{B})^2 - 2 \vec{k} \cdot (\vec{E} \times \vec{B}) | \vec{k} \\
 &= \vec{k}^2 [E^2 \sin^2 \chi(\vec{k}, \vec{E}) + B^2 \sin^2 \chi(\vec{k}, \vec{B}) - 2EB \hat{k} \cdot \vec{s}]
 \end{aligned}$$

with $\hat{k} = \frac{\vec{k}}{|\vec{k}|}$, $E = |\vec{E}|$, $B = |\vec{B}|$, $\vec{s} \equiv \frac{\vec{E} \times \vec{B}}{|\vec{E}||\vec{B}|}$ unit Poynting vector

→ dispersion relations

$$\omega^2(|\vec{k}|) = \vec{k}^2 \left\{ 1 - 2c_i [E^2 \sin^2 \chi(\vec{k}, \vec{E}) + B^2 \sin^2 \chi(\vec{k}, \vec{B}) - 2EB \hat{k} \cdot \vec{s}] \right\} + \alpha \sigma \left(\left(\frac{eF}{m^2} \right)^4 \right)$$

$$\omega = |\vec{k}| \left\{ 1 - c_i [\dots] \right\} + \alpha \sigma \left(\left(\frac{eF}{m^2} \right)^4 \right)$$

(to this order)

$$v_{gr,i} = \frac{d\omega}{d|\vec{k}|} = v_{ph,i} = \frac{\omega}{|\vec{k}|} \equiv v_i \quad \rightarrow \text{index of refraction } n_i = \frac{1}{v_i}$$

$$\rightarrow \left\{ \begin{matrix} v_1 \\ v_2 \end{matrix} \right\} = 1 - \frac{\alpha}{\pi} \frac{1}{g_0} \left\{ \begin{matrix} 7 \\ 4 \end{matrix} \right\} \left[\left(\frac{eE}{m^2} \right)^2 \sin^2 \chi(\vec{k}, \vec{E}) + \left(\frac{eB}{m^2} \right)^2 \sin^2 \chi(\vec{B}, \vec{k}) - 2 \frac{eE}{m^2} \frac{eB}{m^2} \hat{k} \cdot \vec{s} \right] + \alpha \sigma \left(\left(\frac{eF}{m^2} \right)^4 \right)$$

magnetic field

$$\left\{ \begin{matrix} v_1 \\ v_2 \end{matrix} \right\} = 1 - \frac{\alpha}{\pi} \frac{1}{g_0} \left\{ \begin{matrix} 7 \\ 4 \end{matrix} \right\} \left(\frac{eB}{m^2} \right)^2 \sin^2 \chi(\vec{B}, \vec{k}) + \alpha \sigma \left(\left(\frac{eF}{m^2} \right)^4 \right)$$

- $\vec{e}_1 |_{E=0} \sim \omega (\vec{k} \times \vec{B})$ \perp plane spanned by \vec{k}, \vec{B}
- $\vec{e}_2 |_{E=0} \sim -\vec{k} (\vec{k} \cdot \vec{B}) + \omega^2 \vec{B}$ \in plane spanned by \vec{k}, \vec{B}

"1" → \perp mode, "2" → \parallel mode

magnetized quantum vacuum is birefringent [Toll '52]

→ Signal: ellipticity specified by angle $\Delta\Phi := 2\pi \frac{\Delta n}{\lambda} L$
 λ ← probe wavelength, L ← optical path, $\Delta n = n_1 - n_2$