

Non-Extensive Statistical Approach for Hadronization and its Application

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Approach: Eur. Phys. J. A49 (2013) 110, Physica A 392 (2013) 3132

Application: J.Phys.CS 612 (2015) 012048 arXiv:1405.3963, 1501.02352, 1501.05959



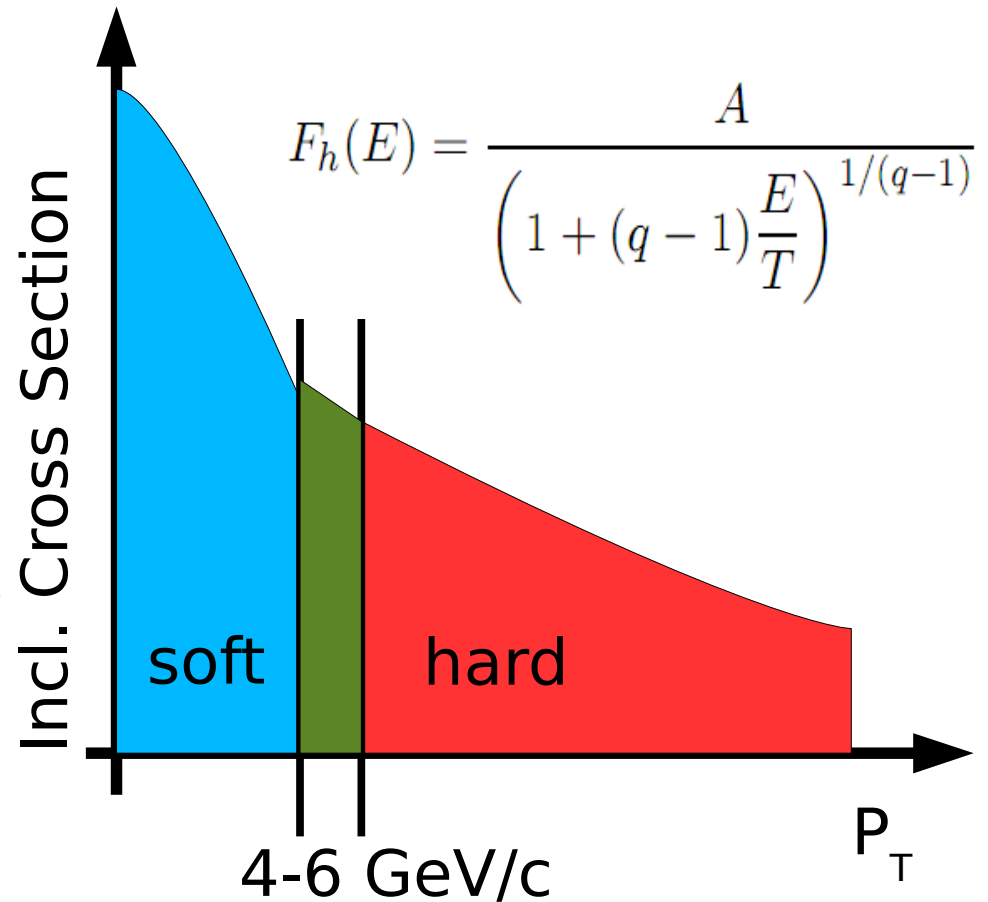
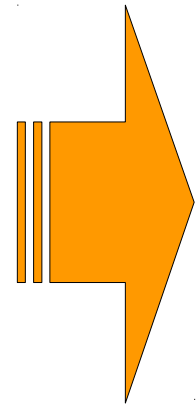
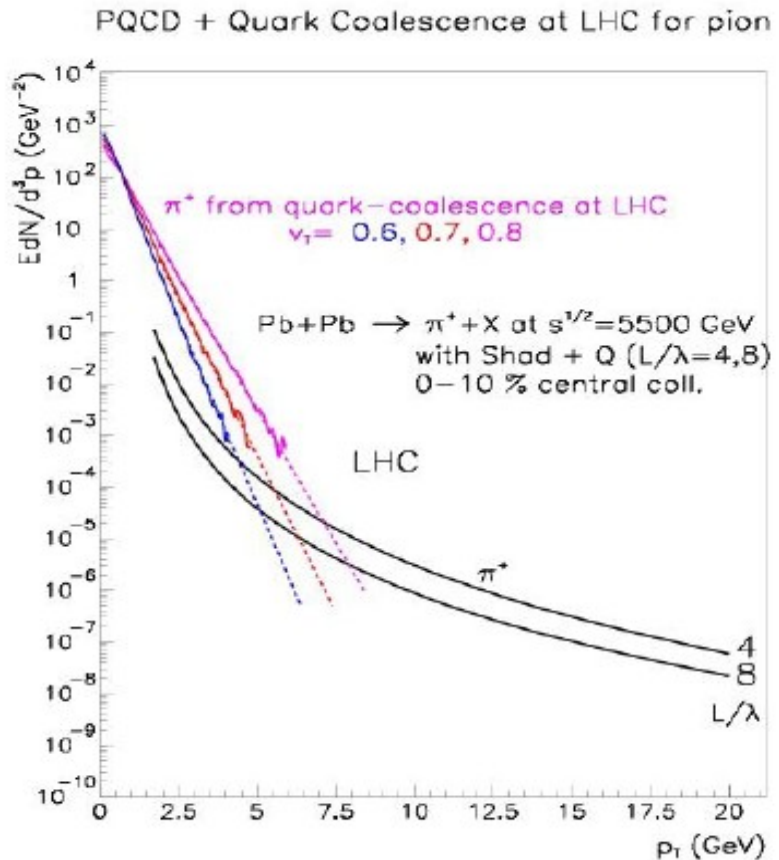
Strangeness in Quark Matter 2015, Dubna, Russia, 10th July 2015

OUTLINE

- Motivation...
 - by a student exercise
- Non-extensive statistical approach
 - Fits of experimental spectra from e^+e^- , pp
 - Non-extensive statistical approach
- Can Tsallis – Pareto fit spectra of HIC?
 - The soft+hard model and its applications
 - Spectra fit and extraction of q and T
 - Asimuthal anisotropy from the model

MOTIVATION

- Simplest and best fit to hadron spectra at low- p_T & high- p_T



P. Lévai, GGB, G. Fai: JPG35, 104111 (2008)

The student exercise...

- Why use Tsallis–Pareto distribution?
 - Is it true Boltzmann-Gibbs fits better at low momenta?
 - Is it true Power-law distribution is better at high momenta?
 - Is it true Tsallis – Pareto fits the whole momentum range?
 - Can we apply this for any system: ee, pp, pA, AA?
- Let's see first a 'known' case:
 - PYTHIA6.4: π , K and p production in proton-proton @ 14 TeV
 - Fits of Boltzmann-Gibbs, Power law, and Tsallis–Pareto distributions
 - Low momenta: [1.2 GeV/c : 2.0 GeV/c] or [1.2 GeV/c : 5.0 GeV/c]
 - High momenta: [5.0 GeV/c : 15.0 GeV/c]
 - Full range: [1.2 GeV/c : 15.0 GeV/c]

What can we learn from a simple exercise?

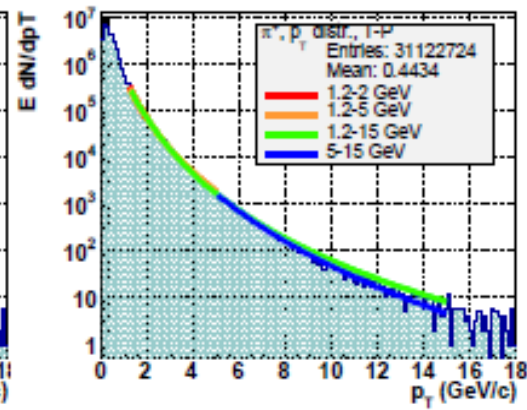
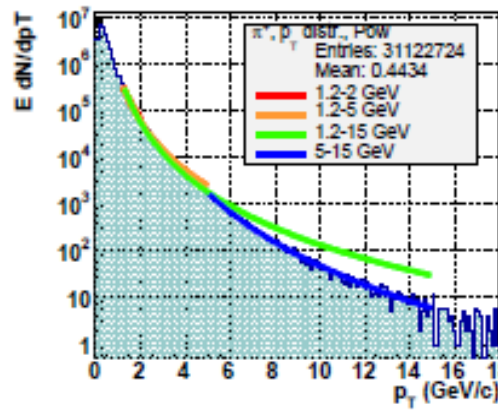
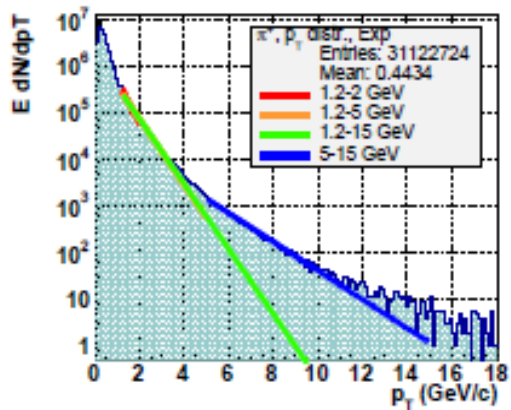
The student exercise...

Boltzmann–Gibbs

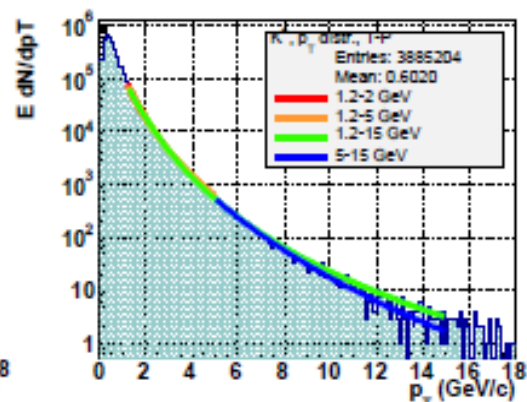
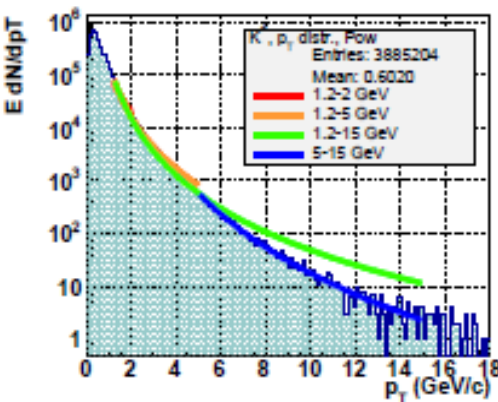
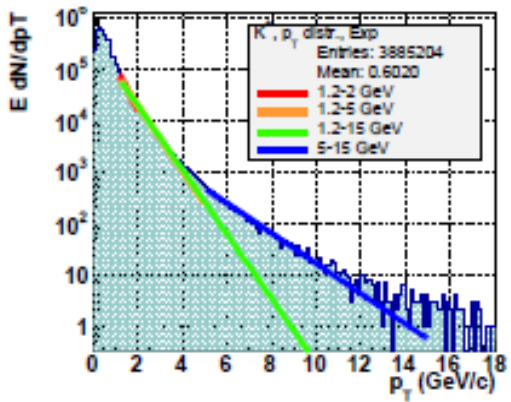
Power Law

Tsallis–Pareto

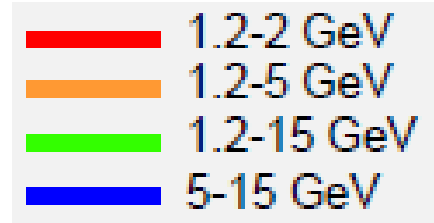
Pions



Kaons



The fitted momentum regions:



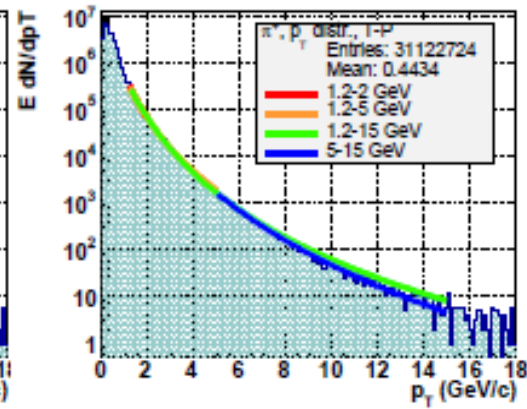
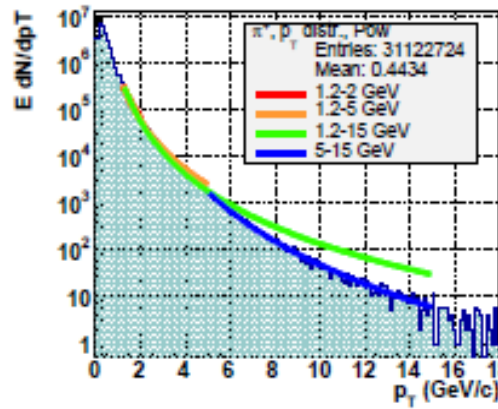
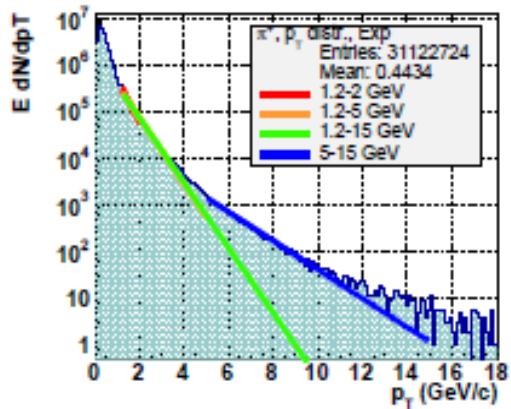
The student exercise...

Boltzmann–Gibbs

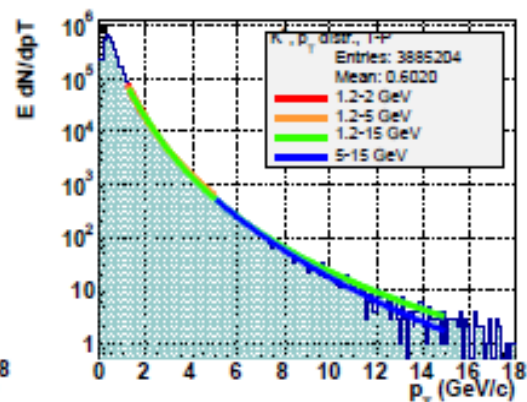
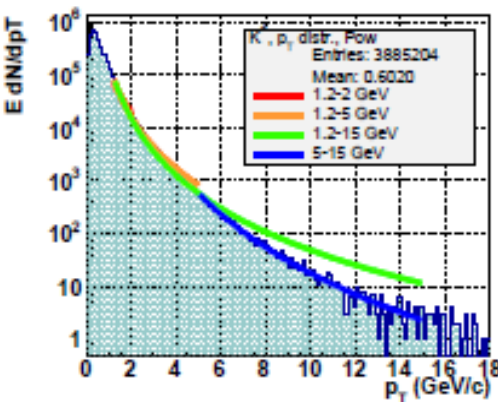
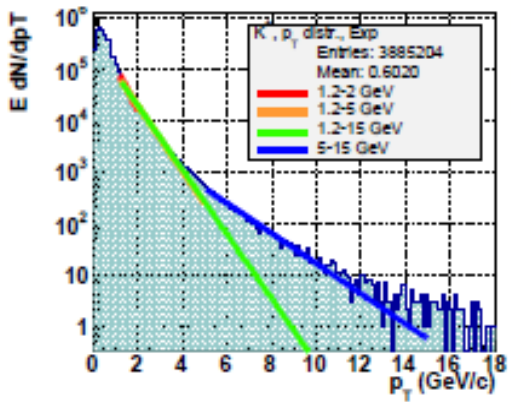
Power Law

Tsallis–Pareto

Pions



Kaons



[1,2:2] GeV/c

[1,2:5] GeV/c

[1,2:15] GeV/c

[5:15] GeV/c

χ^2 values:

Exp	112,37/29,81/27,34	623,89/130,48/109,26	254,12/61,71/48,13	3,01/1,44/1,45
Pow	1,71/0,98/0,47	161,27/55,68/56,08	214,12/76,92/77,26	1,37/1,144/0,91
TP	0,45/1,19/0,56	12,21/5,55/11,06	10,39/4,37/7,77	1,14/0,97/0,91

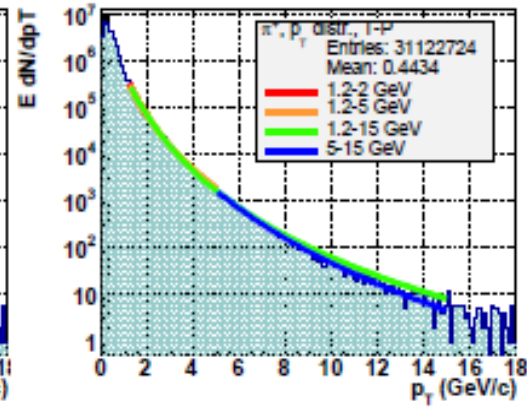
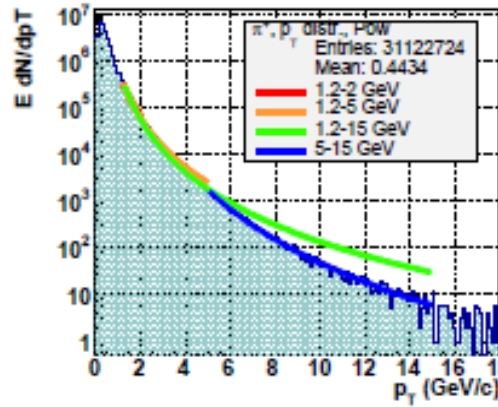
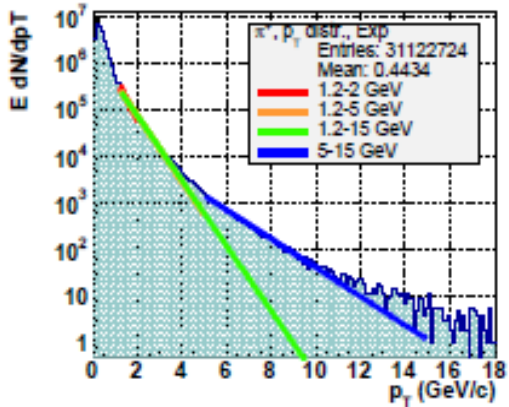
The student exercise...

Boltzmann–Gibbs

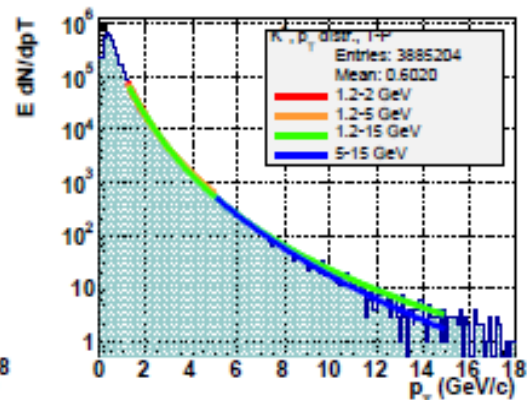
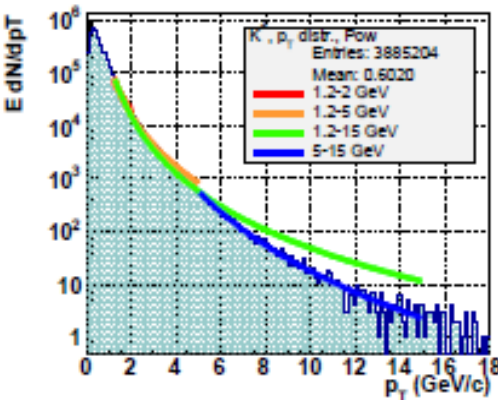
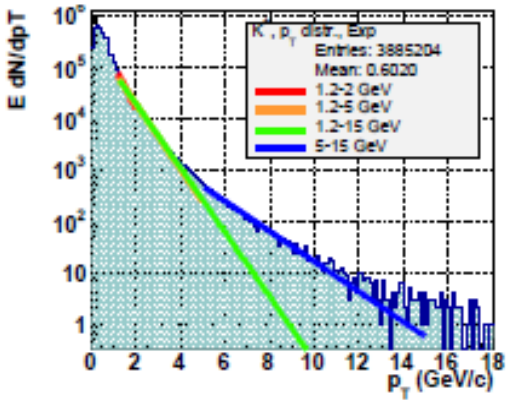
Power Law

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Kaons



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χ^2 values:

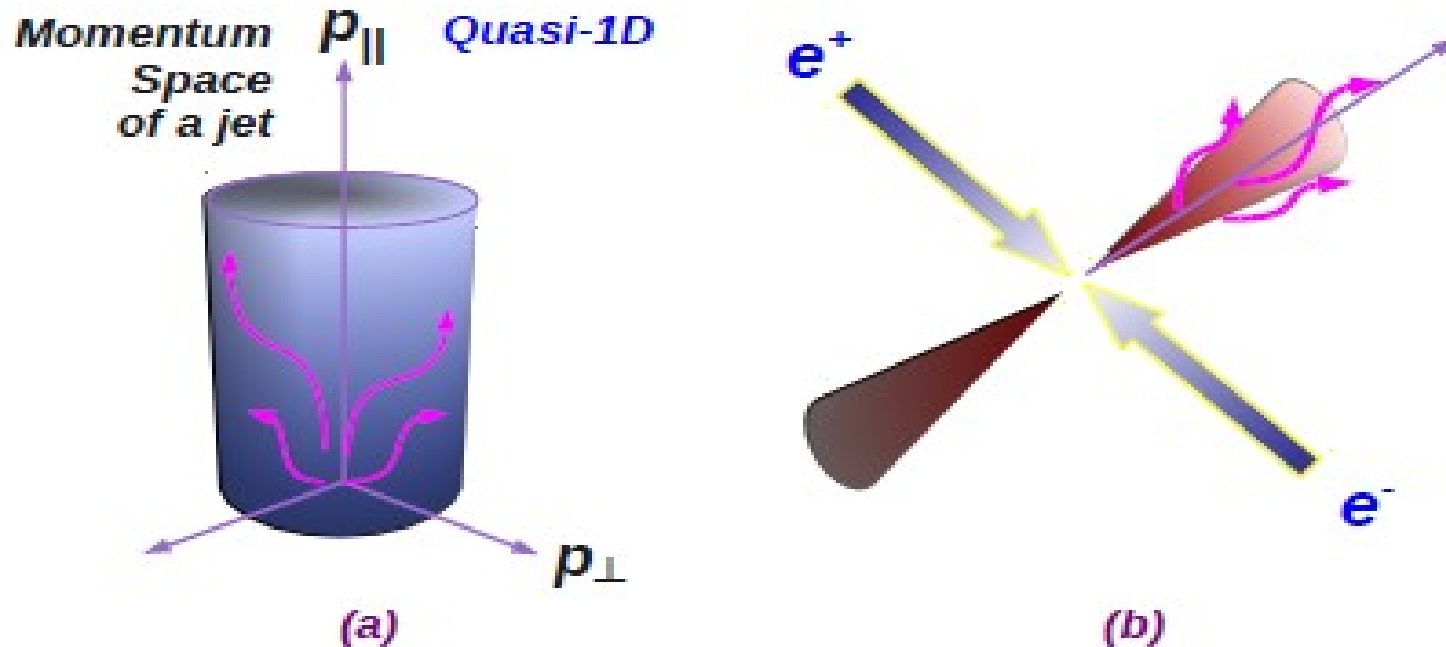


The student exercise...

- Why fit Tsallis–Pareto distribution?
 - Yes, it is true Boltzmann-Gibbs fits better at low momenta.
 - Yes, it is true Power-law distribution is better at high momenta.
 - Yes, it is true Tsallis – Pareto fits the whole momentum range.
 - Can we apply this for any system: ee, pp, pA, AA?
- But carefully
 - BODY vs. TAIL (dependence on the momentum regions)
 - Need to find the proper variable E_{jet} , p_T , m_T , m_T^*
 - Need for
 - High- p_T PID hadron data
 - High statistic data
 - Spectra in several multiplicity bins
 - Dream: all of these on track-by-track basis

Application of the non-extensive statistical approach on small systems using experimental data.

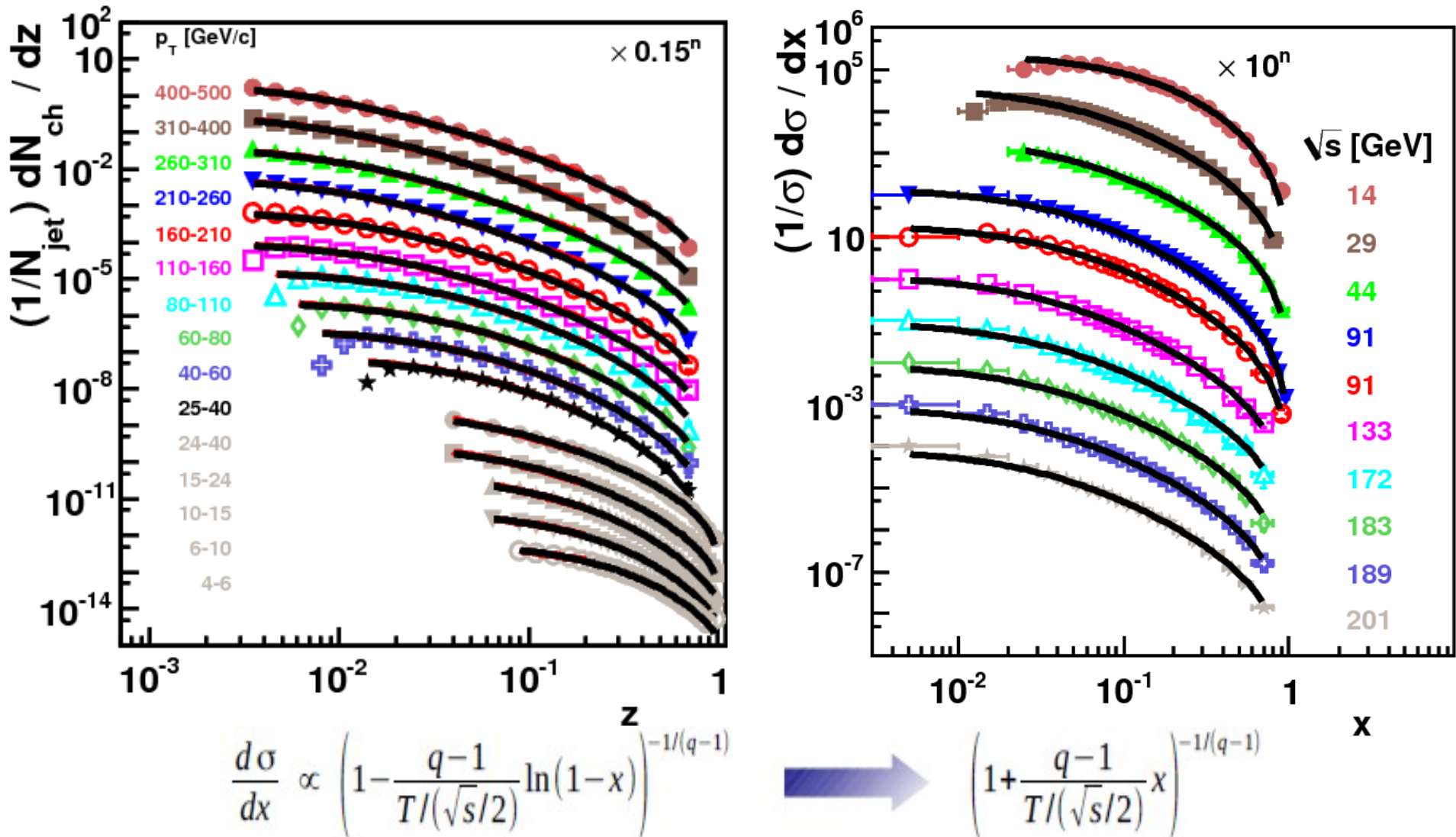
The 'Thermodynamics of Jets'



K. Ürmössy, G.G. Barnaföldi, T.S. Bíró:

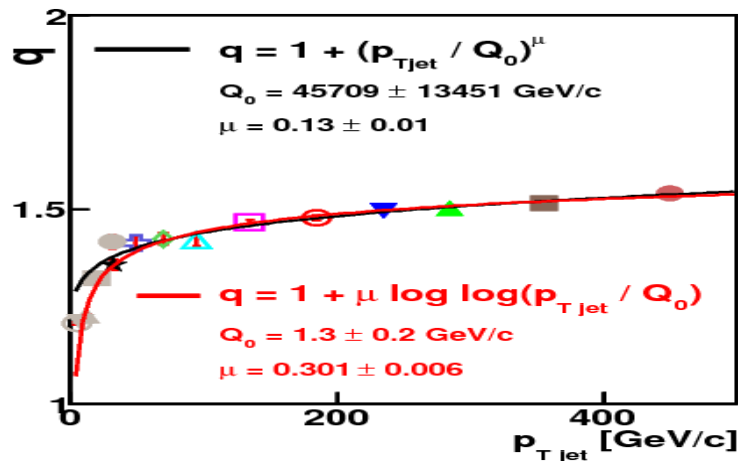
- Microcanonical Jet-Fragmentation in pp at LHC energies:
Phys. Lett. B701 (2011) 111
- Generalized Tsallis distribution in e^+e^- collisions
Phys. Lett. B718 (2012) 125

Fits for jet spectra in pp (left) and e⁺e⁻ (right)

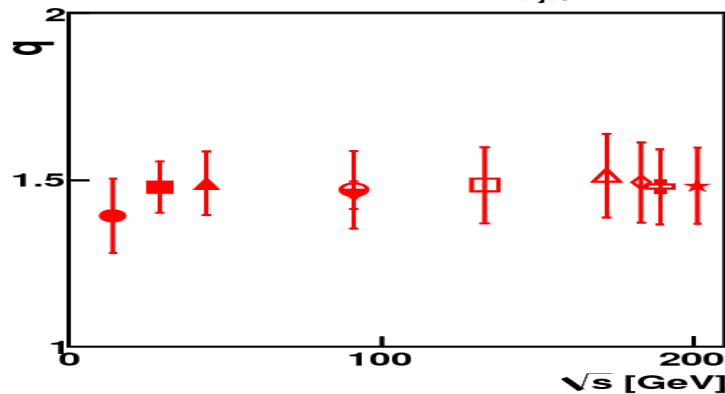
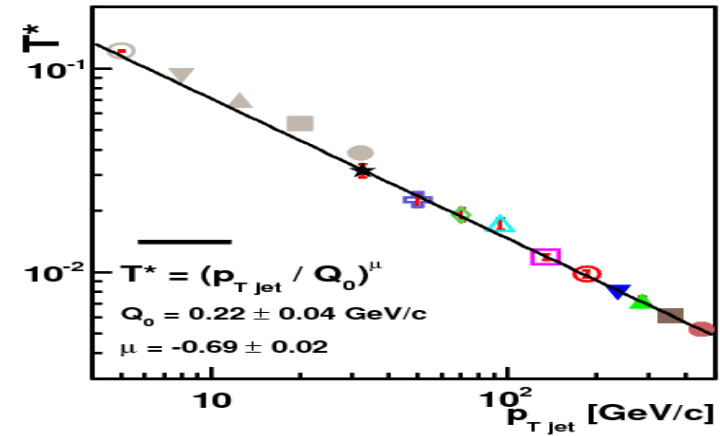


Ref: K Ürmössy, GGB, TS Biró, PLB 710 (2011) 111, PLB 718 (2012) 125.

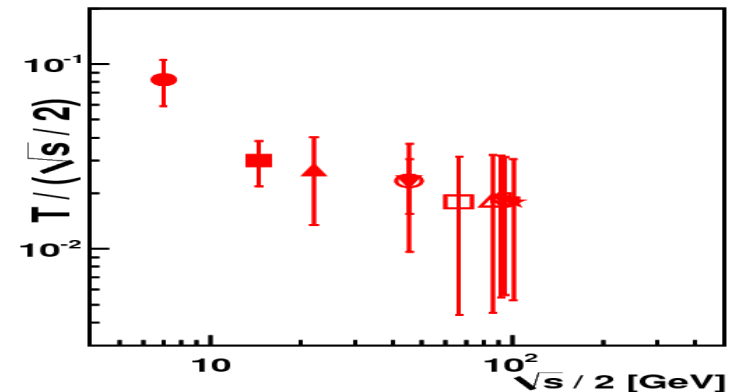
The evolution of q and T parameters



pp



e^+e^-



K Ürmössy, GGB, TS Biró,
 PLB 710 (2011) 111, PLB 718 (2012) 125.

- Energy dependence (hard)

- Parameters q seem to saturate at high energies $q > 1.1$
- Parameter T is decreasing with increasing energy

What is the physical meaning of these
' q ' and ' T ' parameters?

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The non-extensive statistical approach

- Extensive Boltzmann – Gibbs statistics

$$\begin{aligned} S_{12} &= S_1 + \hat{S}_2 \\ E_{12} &= E_2 + E_2 \end{aligned} \quad \rightarrow \quad S_B = - \sum_i p_i \ln p_i$$



The non-extensive statistical approach

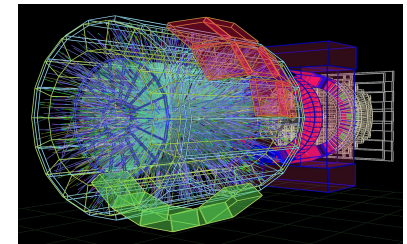
- Extensive Boltzmann – Gibbs statistics

$$\begin{aligned} S_{12} &= S_1 + S_2 & \longrightarrow & S_B = - \sum_i p_i \ln p_i \\ E_{12} &= E_1 + E_2 \end{aligned}$$



- Non-extensivity → generalized entropy

$$\begin{aligned} \hat{L}_{12}(S_{12}) &= \hat{L}_1(S_1) + \hat{L}_2(S_2) & \longrightarrow & S_T = \frac{1}{1-q} \sum_i (p_i^q - p_i) \\ L_{12}(E_{12}) &= L_1(E_1) + L_2(E_2) \end{aligned}$$

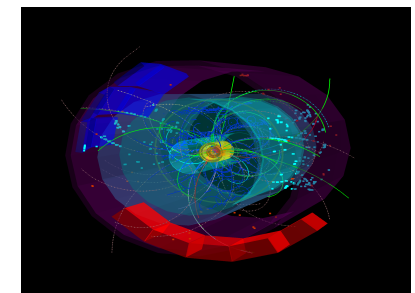


- Tsallis entropy

$$S_{12} = S_1 + S_2 + (q-1)S_1S_2 \longrightarrow \hat{L}(S) = \frac{1}{q-1} \ln(1 + (q-1)S)$$

from here: Tsallis – Pareto distribution

$$f(\varepsilon) = \left[1 + (q-1) \frac{\varepsilon}{T} \right]^{-\frac{1}{q-1}}$$



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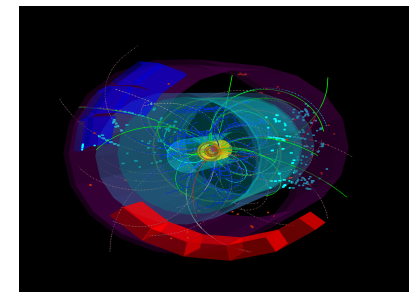
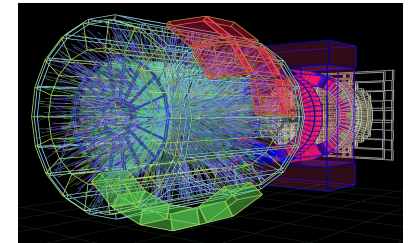
The non-extensive statistical approach

- Tsallis – Pareto distribution

$$f(\varepsilon) = \left[1 + (q-1) \frac{\varepsilon}{T} \right]^{-\frac{1}{q-1}}$$

$$q = \frac{\langle S'(E)^2 + S''(E) \rangle}{\langle S'(E) \rangle^2}$$

$$\frac{1}{T} = \langle S'(E) \rangle$$



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The non-extensive statistical approach

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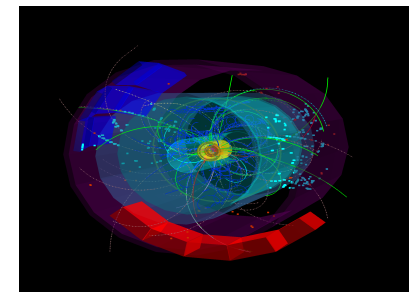
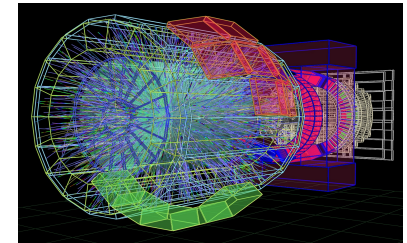
$$\frac{1}{T} = \langle S'(E) \rangle$$

$$q = \frac{\langle n(n-1) \rangle}{\langle n \rangle^2}$$

$$q = 1 + \frac{\Delta T^2}{T^2} - \frac{1}{C}$$

$$T = \frac{E}{\langle n \rangle}$$

$$T = \frac{\int \varepsilon f_{TS}(\varepsilon)}{\int f_{TS}(\varepsilon)} = \frac{DT}{1 - (q-1)(D+1)}$$



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The non-extensive statistical approach

Hadron spectra in pp collisions can be described by the *Tsallis distribution*:

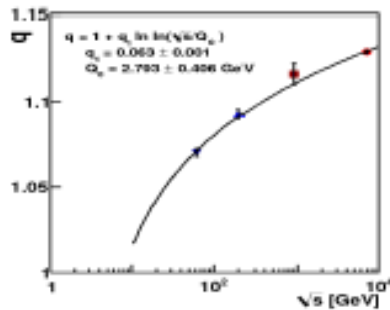
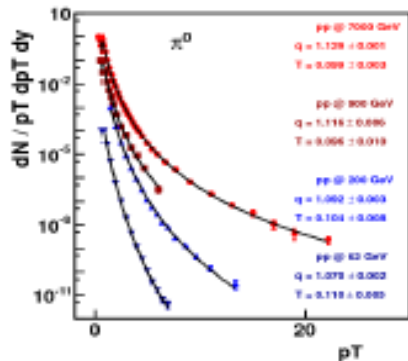
$$\frac{dN}{d^3 p} \propto \left[1 + \frac{q-1}{T} (m_T - m) \right]^{-1/(q-1)}$$

π spectra in pp collisions depends similarly on \sqrt{s} and on the multiplicity N

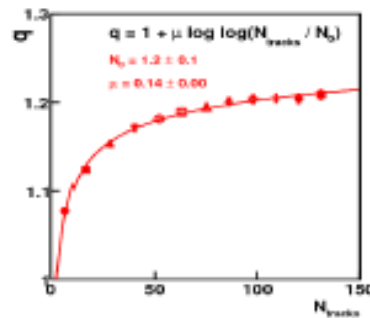
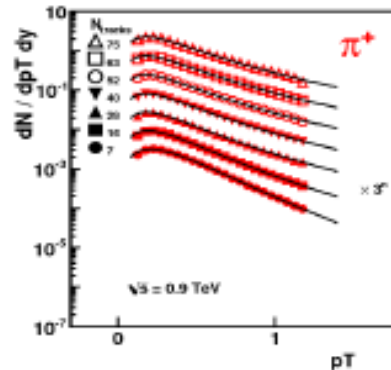
$$q(s) = 1 + q_1 \ln \ln(\sqrt{s}/Q_0),$$

$$q(N) = 1 + \mu \ln \ln(N/N_0).$$

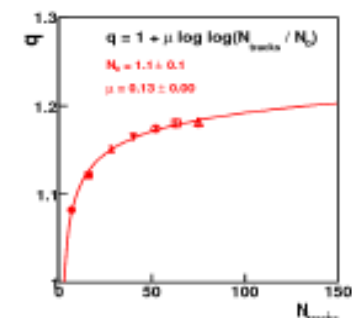
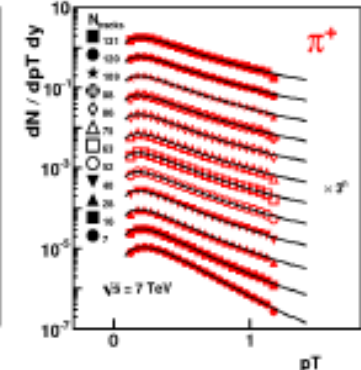
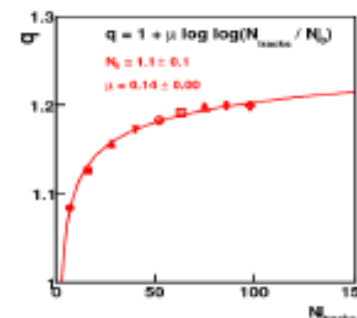
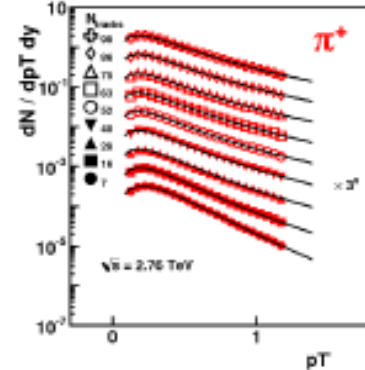
$\sqrt{s} = \text{fix}$



$N = \text{fix}$



arXiv:1405.3963, 1501.02352, 1501.05959



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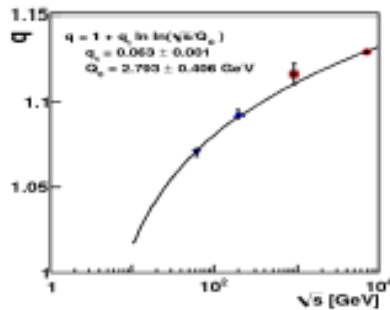
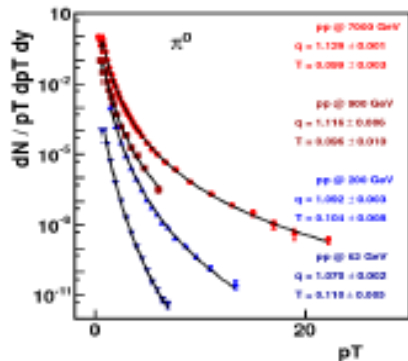
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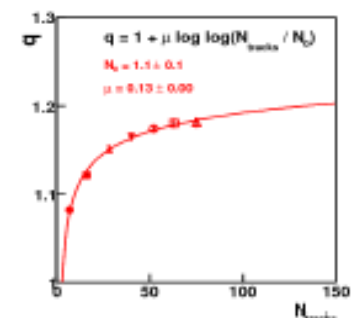
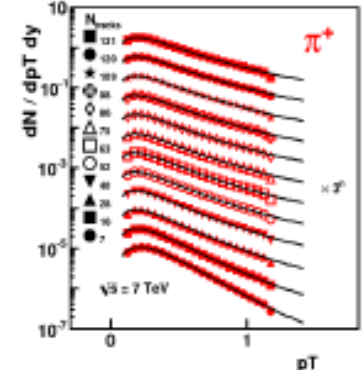
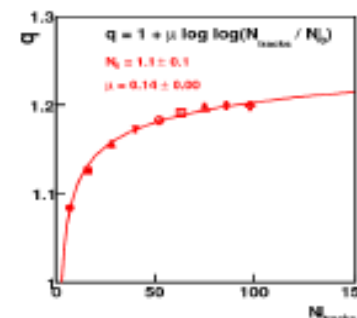
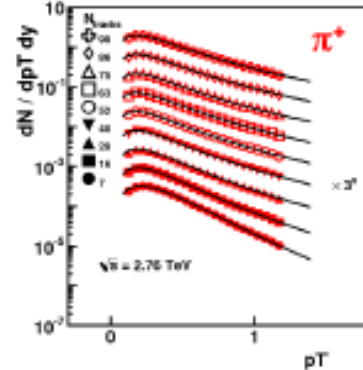
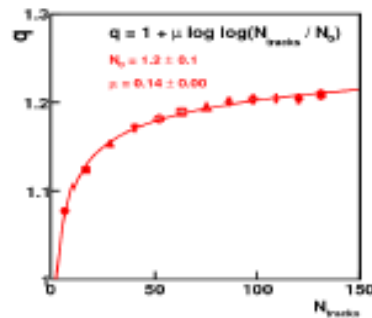
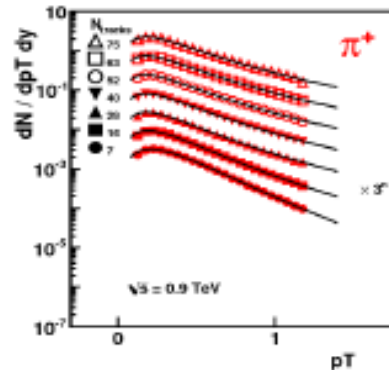
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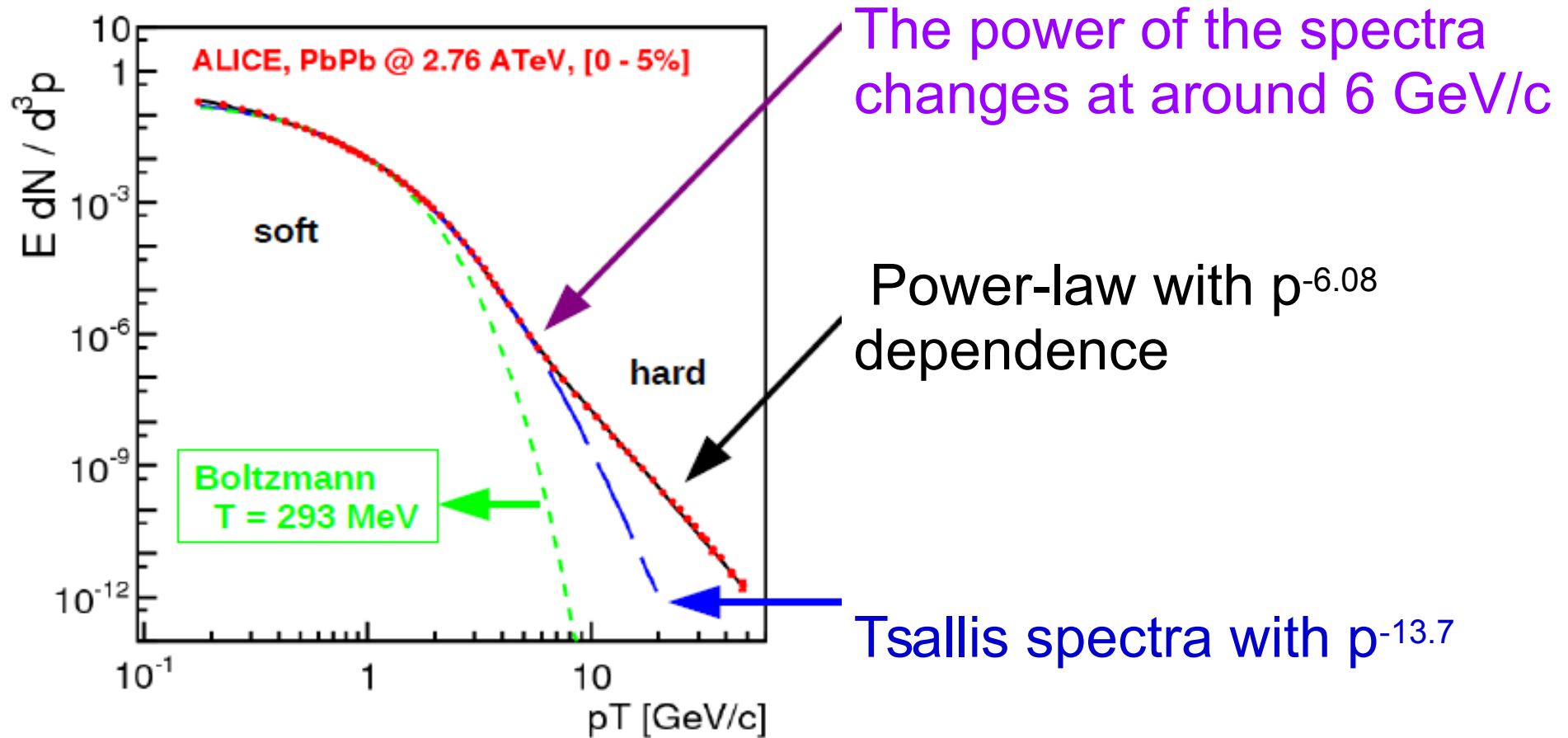
$N = \text{fix}$



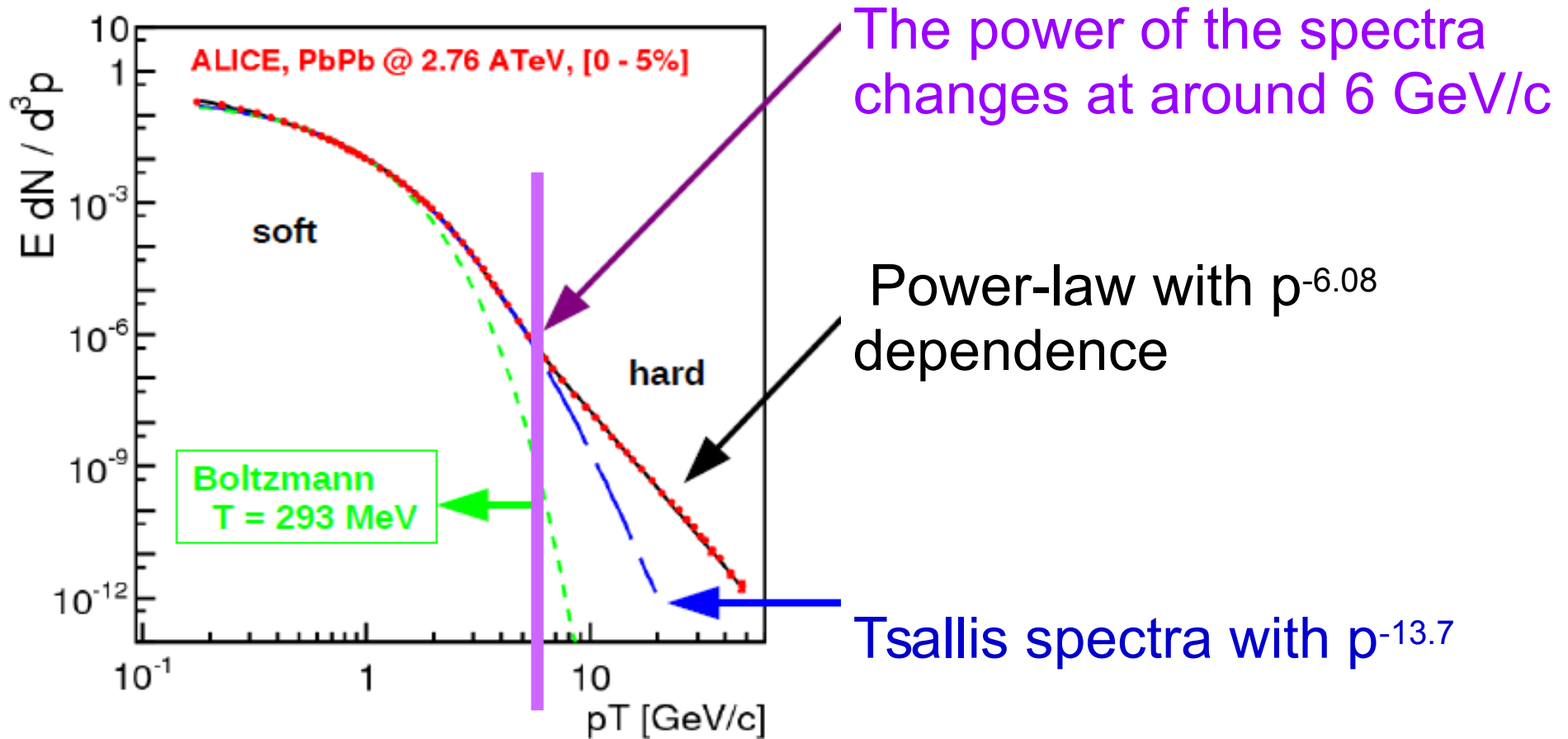
arXiv:1405.3963, 1501.02352, 1501.05959

What if, we would apply this for
a bigger system (AA)
where
Boltzmann–Gibbs
use to work?

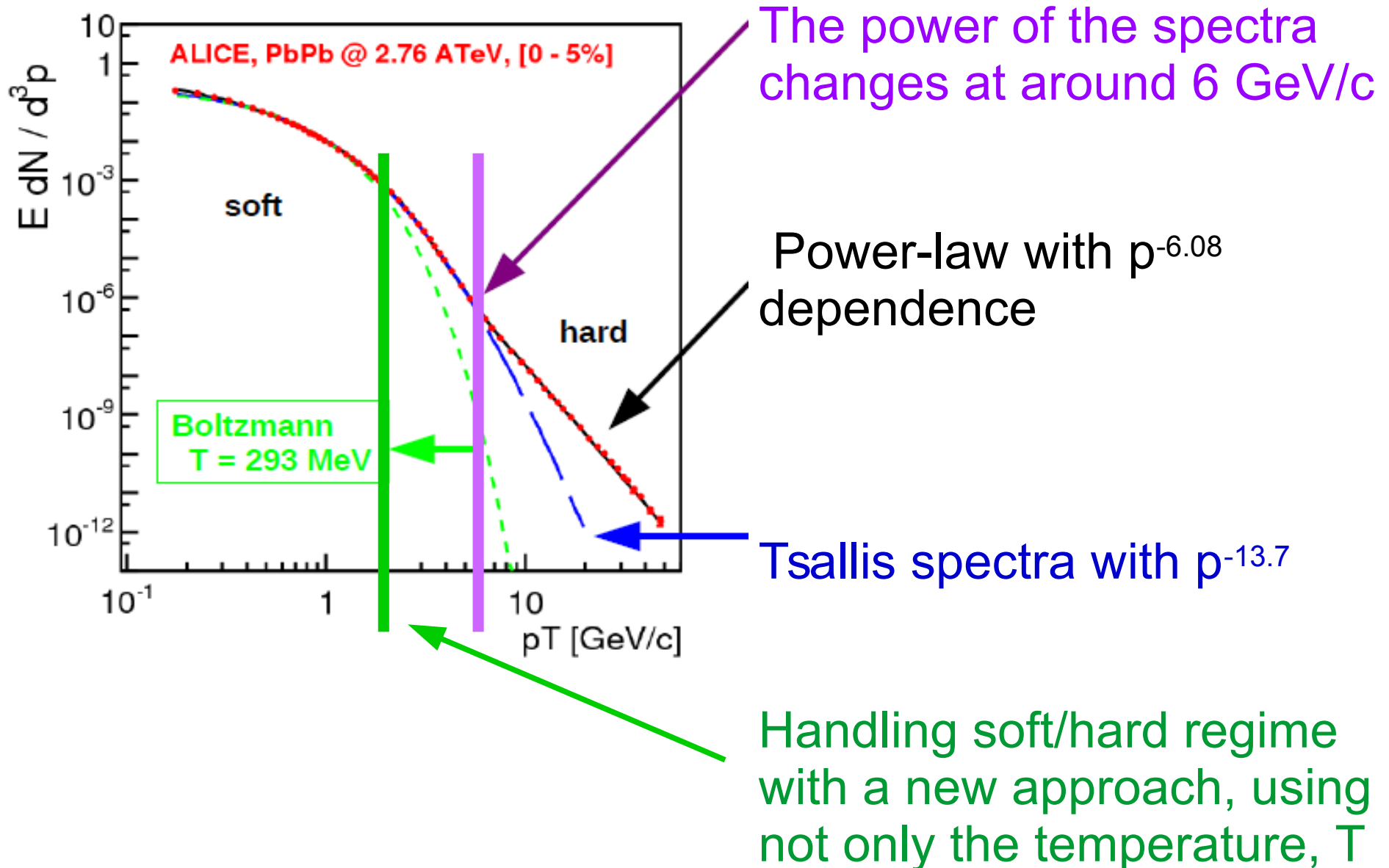
Test with real data in PbPb



Test with real data in PbPb



Test with real data in PbPb



The soft + hard model

- Simplest approximation: soft ('bulk') + hard ('jet') contribution

$$p^0 \frac{dN}{d^3\mathbf{p}} = p^0 \frac{dN^{\text{hard}}}{d^3\mathbf{p}} + p^0 \frac{dN^{\text{soft}}}{d^3\mathbf{p}}$$

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The soft + hard model

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- Identified hadron spectra is given by double Tsallis–Pareto:

$$\left. \frac{dN}{2\pi p_T dp_T dy} \right|_{y=0} = f_{\text{hard}} + f_{\text{soft}} \quad f_i = A_i \left[1 + \frac{(q_i - 1)}{T_i} [\gamma_i(m_T - v_i p_T) - m] \right]^{-1/(q_i - 1)}$$

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in where parameters are given by

- Lorentz factor $\gamma_i = 1/\sqrt{1 - v_i^2}$
- Transverse mass $m_T = \sqrt{p_T^2 + m^2}$
- Doppler temperature $T_i^{\text{Dopp}} = T_i \sqrt{\frac{1 + v_i}{1 - v_i}}$

- Finally we assume N_{part} scaling for the parameters

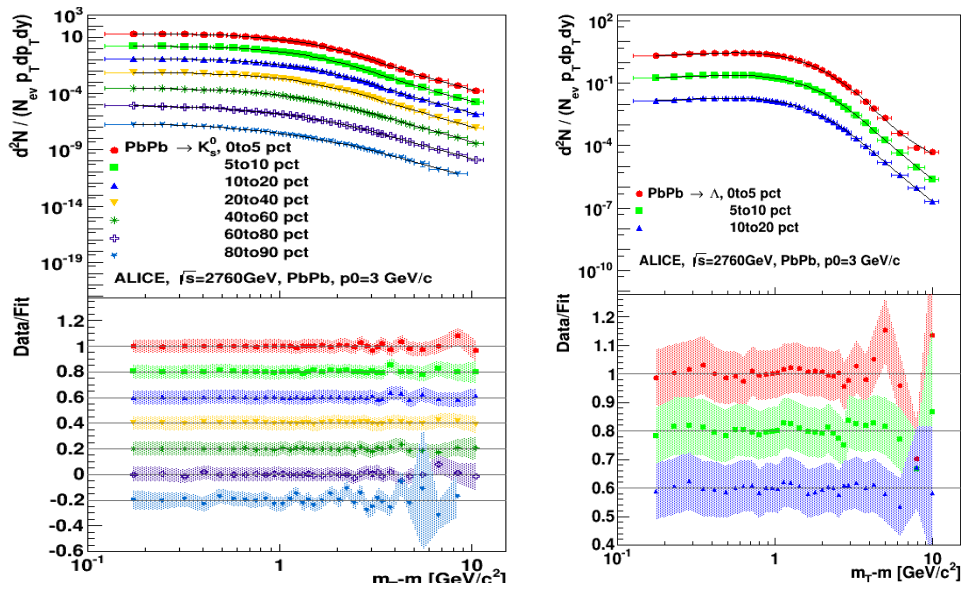
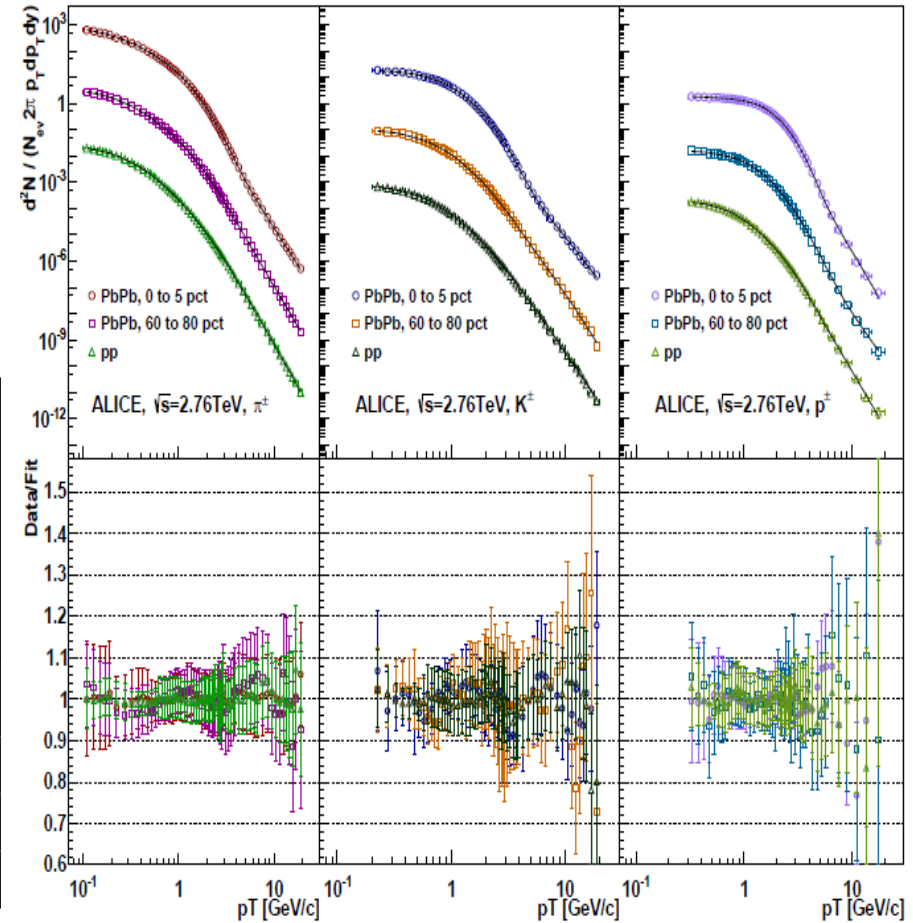
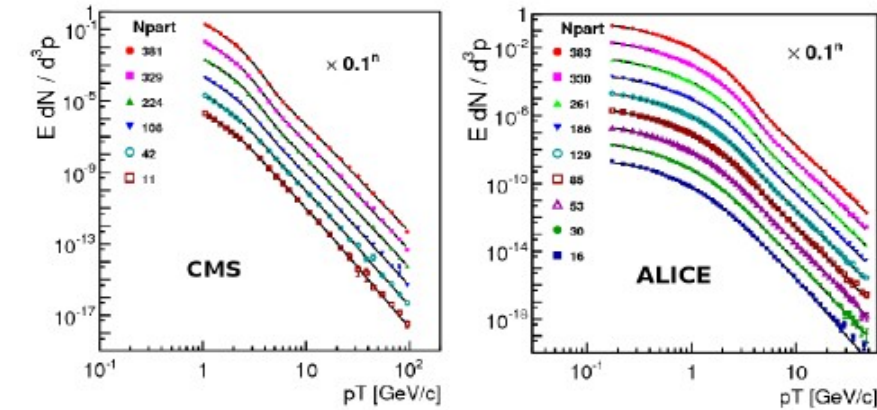
$$\begin{aligned} q_i &= q_{2,i} + \mu_i \ln(N_{\text{part}}/2) \\ T_i^{\text{Dopp}} &= T_{1,i} + \tau_i \ln(N_{\text{part}}) . \end{aligned}$$

arXiv:1405.3963, 1501.02352, 1501.05959
J.Phys.CS 612 (2015) 012048

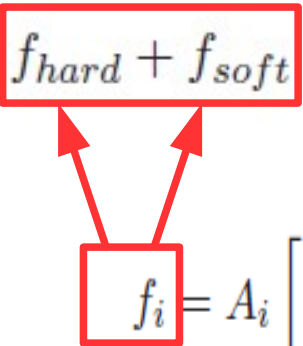
Fit of pp and PbPb (centra/peripheral) data

arXiv:1405.3963, 1501.02352, 1501.05959

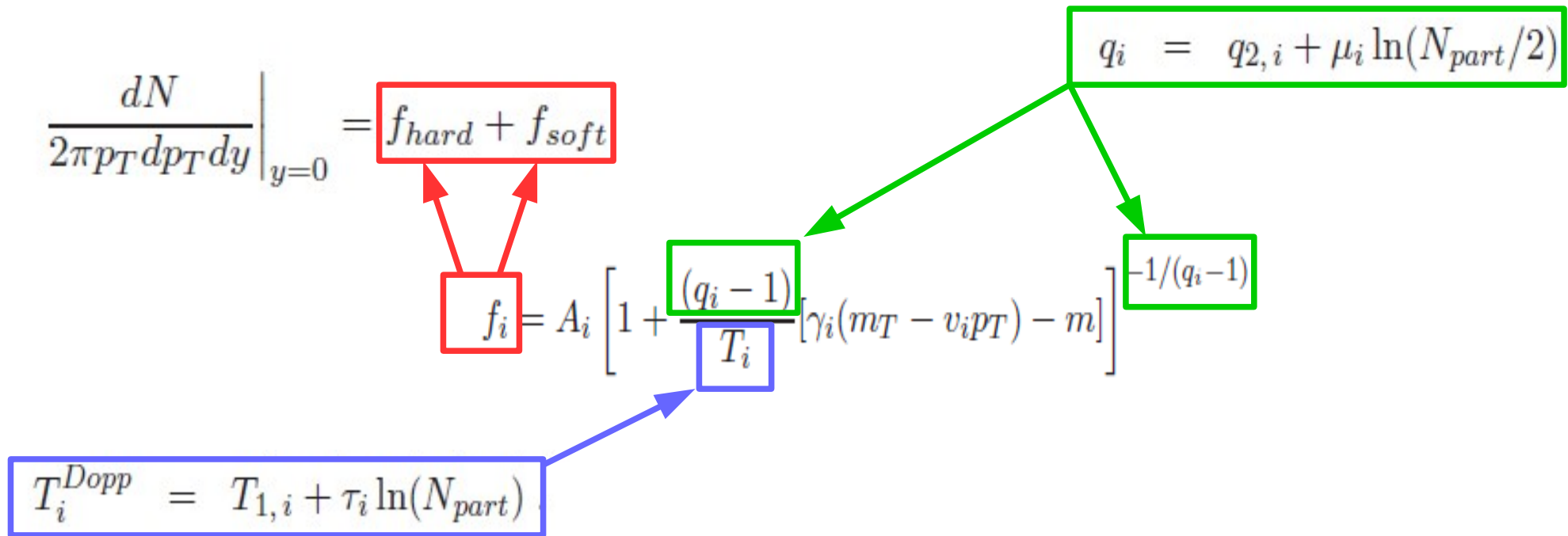
J.Phys.CS 612 (2015) 012048



Parameters of the soft+hard model

$$\frac{dN}{2\pi p_T dp_T dy} \Big|_{y=0} = f_{hard} + f_{soft}$$

$$f_i = A_i \left[1 + \frac{(q_i - 1)}{T_i} [\gamma_i(m_T - v_i p_T) - m] \right]^{-1/(q_i - 1)}$$

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$$q_i = q_{2,i} + \mu_i \ln(N_{part}/2)$$

$$T_i^{Dopp} = T_{1,i} + \tau_i \ln(N_{part})$$

	$q_{2,soft}$	$q_{2,hard}$	μ_{soft}	μ_{hard}
CMS	1.058 ± 0.025	1.136 ± 0.001	-0.008 ± 0.005	0.005 ± 0.0003
ALICE	1.074 ± 0.018	1.131 ± 0.002	-0.009 ± 0.004	0.006 ± 0.0006
PHENIX	1.073 ± 0.016	1.100 ± 0.002	-0.005 ± 0.004	0.000 ± 0.0006

	T_1^{soft} [MeV]	T_1^{hard} [MeV]	τ_{soft} [MeV]	τ_{hard} [MeV]
CMS	310 ± 20	126 ± 5	9.9 ± 3.7	5.3 ± 0.8
ALICE	266 ± 16	194 ± 2	11.5 ± 2.9	-12.5 ± 0.5
PHENIX	165 ± 26	192 ± 20	9.3 ± 5.5	18.7 ± 4.6

The N_{part} scaling of the q & T parameters

- Scaling of the $q_i = q_{2,i} + \mu_i \ln(N_{part}/2)$

- Soft component, $q \rightarrow 1$

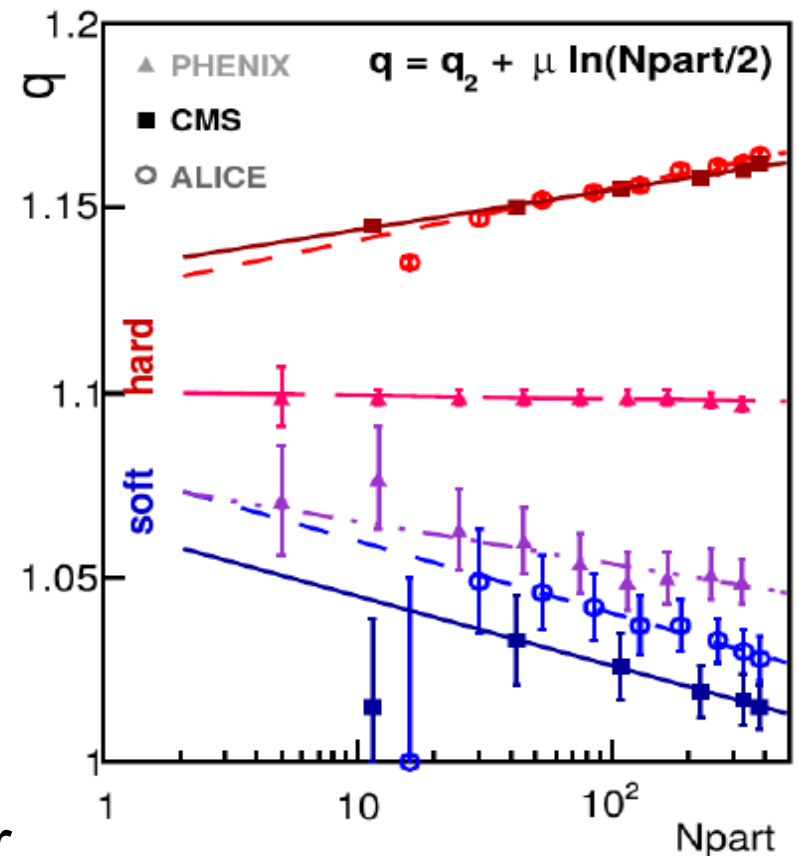
- LHC: decreasing
- RHIC: decreasing

Higher N_{part} result BG statistics

- Hard component, $q > 1.1$

- LHC: slight increasing
- RHIC: constant

Without the soft part result clearer non-extensive behaviour, like e^+e^-



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The N_{part} scaling of the q & T parameters

- Scaling of the $T_i^{Dopp} = T_{1,i} + \tau_i \ln(N_{part})$

- Soft component, $T \sim 200-400$ MeV

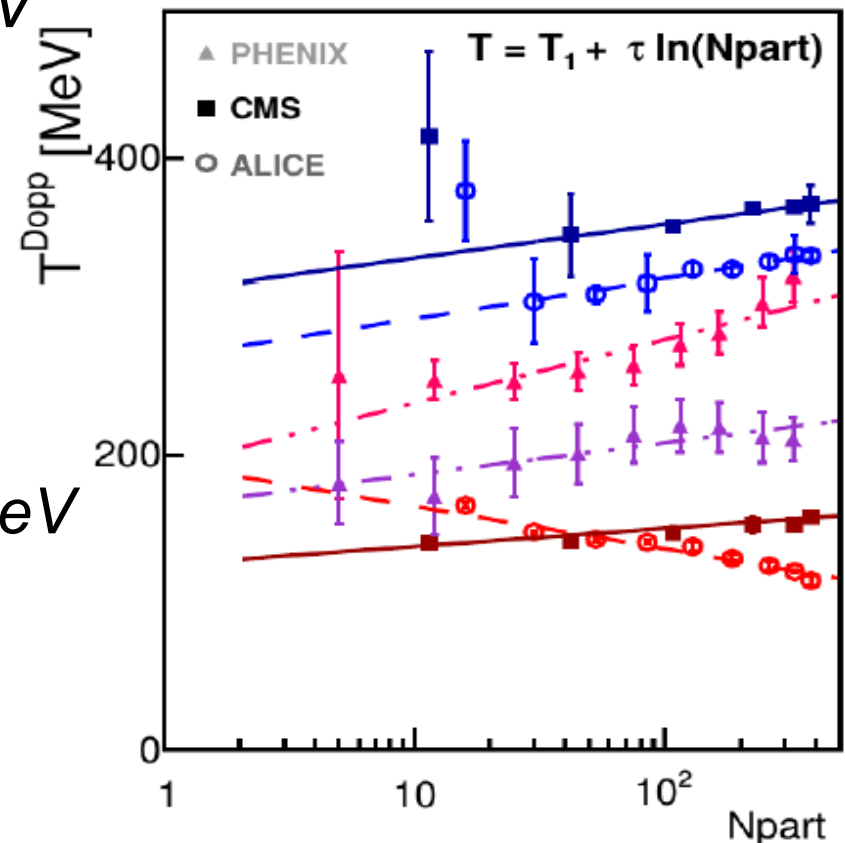
- LHC: constant/increasing
- RHIC: slightly increasing

higher N_{part} results bit higher T

- Hard component, $T \sim 100-300$ MeV

- LHC: decreasing
- RHIC: increasing

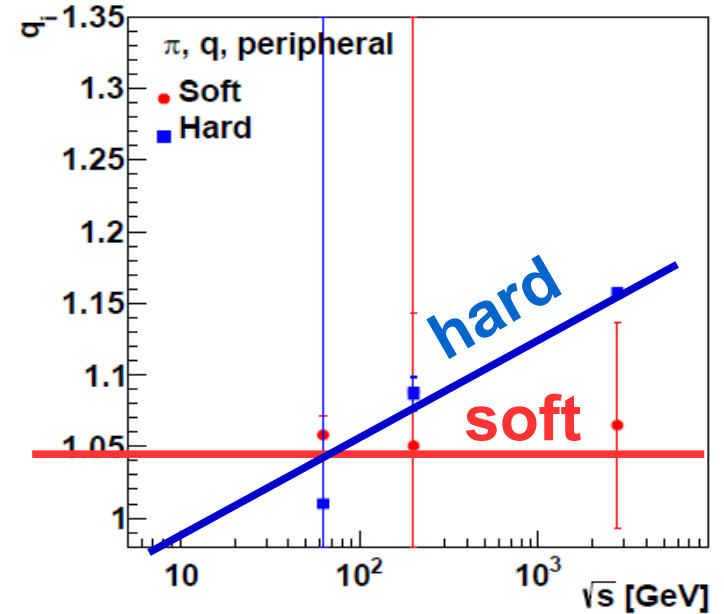
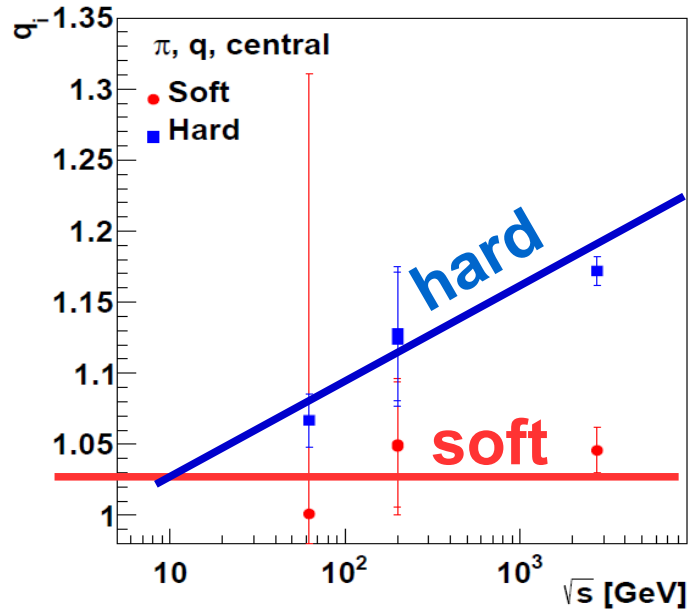
N_{part} scaling seems sensitive...



arXiv:1405.3963, 1501.02352, 1501.05959
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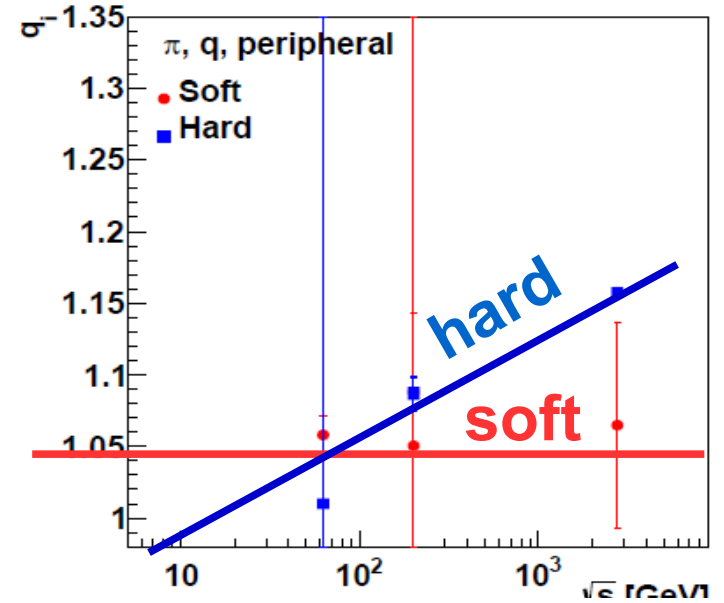
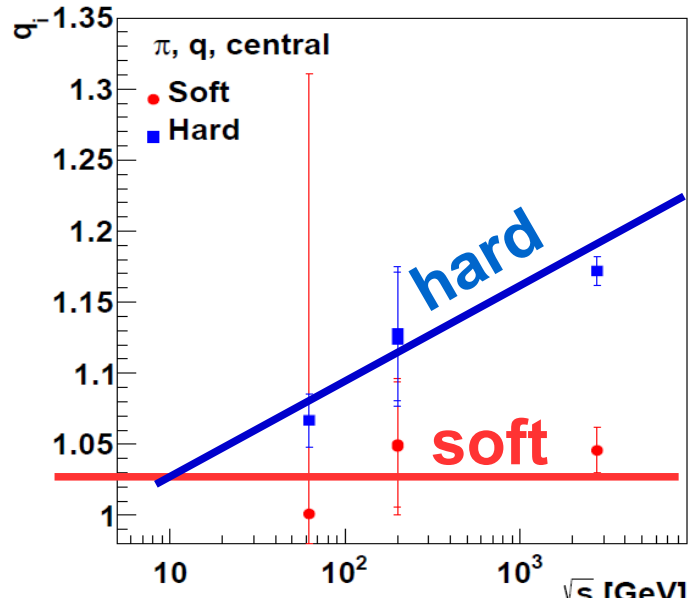
The c.m. energy dependence of q & T

q measures
non-
extensivity

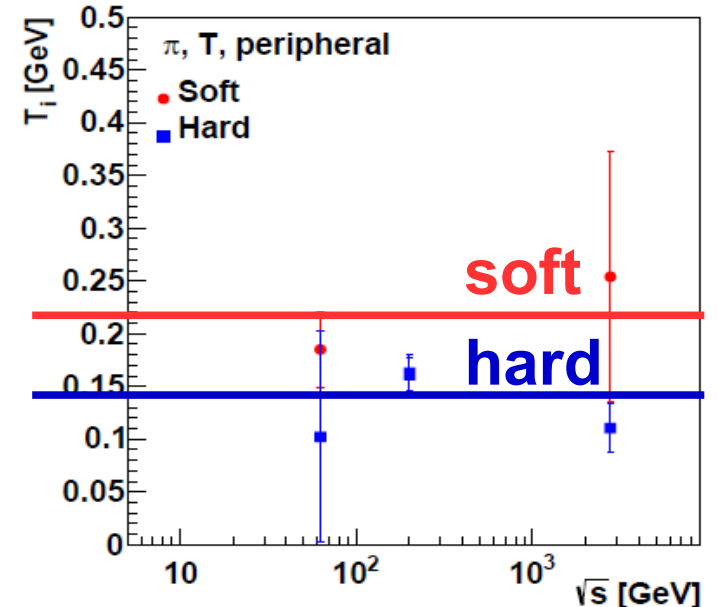
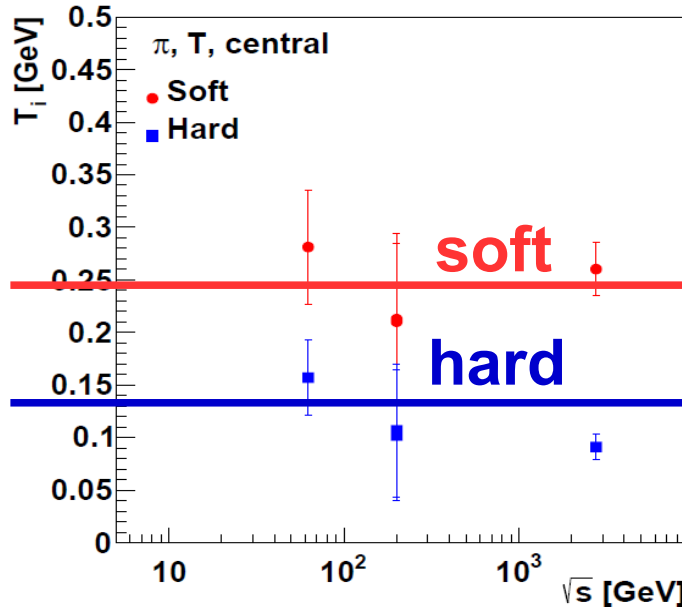


The c.m. energy dependence of q & T

q measures non-extensivity



T measures average E per multiplicity



The c.m. energy dependence of q & T

- Energy dependence

- Parameter q

- **HARD:** clearly increasing
- **SOFT:** no relevant change

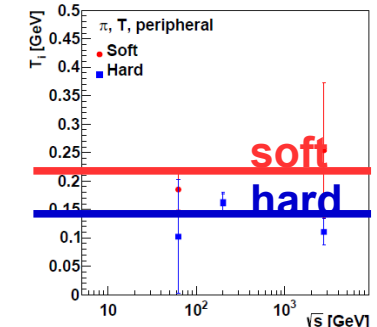
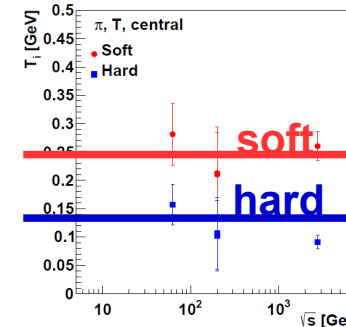
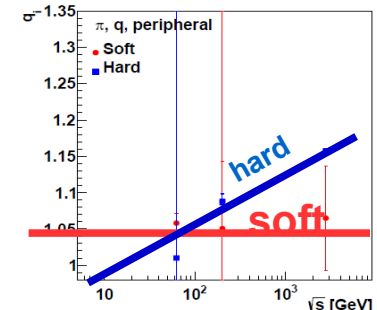
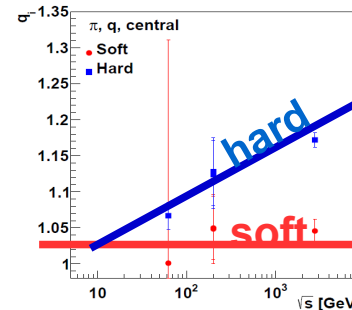
- Parameter T

- **HARD:** central decreasing
peripheral const?

$$T_{centr} = T_{periph}$$

- **SOFT:** similar trend

$$T_{centr} \sim 100 \text{ MeV higher}$$



The c.m. energy dependence of q & T

- Energy dependence

- Parameter q

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- **SOFT:** no relevant change

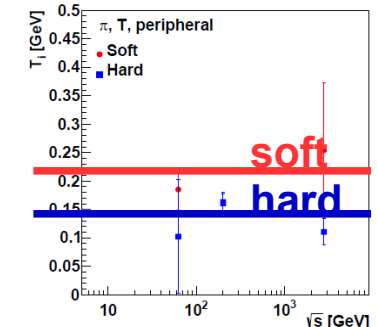
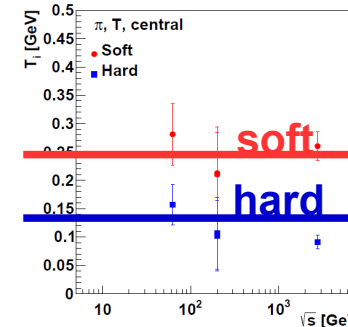
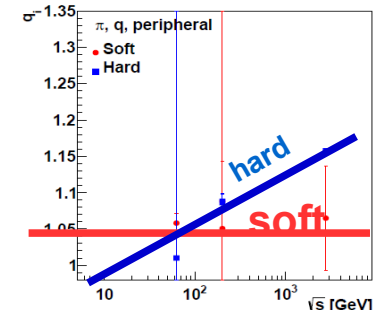
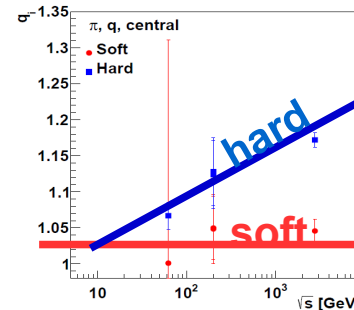
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- Energy dependence

- Parameters q & T present different values for centr./periph.
- Above RHIC soft is BG-like and hard is more TP-like.

Can we connect this to
asimuthal anisotropy?

Connecting spectra and v_2

- Spectra originating from hadronic sources

$$p^0 \frac{dN}{d^3p} \Big|_{y=0} = \int_{-\infty}^{+\infty} d\zeta \int_0^{2\pi} d\alpha f[u_\mu p^\mu] \longrightarrow \frac{dN}{2\pi p_T dp_T dy} \Big|_{y=0} = \int_0^{2\pi} \frac{d\varphi}{2\pi} p^0 \frac{dN}{d^3p} \Big|_{y=0}$$

where we used parameters and assumptions

- Hadron momentum: $p^\mu = (m_T \cosh y, m_T \sinh y, p_T \cos \varphi, p_T \sin \varphi)$

- Cylindric symmetry: $u^\mu = (\gamma \cosh \zeta, \gamma \sinh \zeta, \gamma v \cos \alpha, \gamma v \sin \alpha)$

where $\zeta = \frac{1}{2} \ln[(t+z)/(t-z)]$ and $\gamma = 1/\sqrt{1-v^2}$,

- Co-moving energy: $u_\mu p^\mu \Big|_{y=0} = \gamma [m_T \cosh \zeta - v p_T \cos(\varphi - \alpha)]$

- Transverse flow: $v(\alpha) = v_0 + \sum_{m=1}^{\infty} \delta v_m \cos(m\alpha) \equiv v_0 + \delta v(\alpha)$

- Taylor expansion: $f[u_\mu p^\mu] \Big|_{y=0} = \sum_{m=0}^{\infty} \frac{[\delta v(\alpha)]^m}{m!} \frac{\partial^m}{\partial v_0^m} f[u_\mu p^\mu] \Big|_{y=0}^{v(\alpha)=v_0}$

Connecting spectra and v_2

- Spectra originating from hadronic sources

$$\left. \frac{dN}{2\pi p_T dp_T dy} \right|_{y=0} = \int_0^{2\pi} \frac{d\varphi}{2\pi} p^0 \left. \frac{dN}{d^3 p} \right|_{y=0} = \sum_{m=0}^{\infty} \frac{a_m}{m!} \frac{\partial^m}{\partial v_0^m} f[E(v_0)] \approx f[E(v_0)] + O(\delta v^2)$$

where $E(v_0) = \gamma_0(m_T - v_0 p_T)$ and $a_m = \int_0^{2\pi} d\alpha [\delta v(\alpha)]^m$.

- Azimuthal anisotropy:

$$v_n = \frac{\int_0^{2\pi} d\varphi \cos(n\varphi) p^0 \left. \frac{dN}{d^3 p} \right|_{y=0}}{\int_0^{2\pi} d\varphi p^0 \left. \frac{dN}{d^3 p} \right|_{y=0}} \approx \frac{\delta v_n \gamma_0^3 (v_0 m_T - p_T) f'[E(v_0)]}{2 f[E(v_0)]} + O(\delta v^2)$$

– Boltzmann–Gibbs: $\longrightarrow v_n^{\text{BG}} \approx \frac{\delta v_n \beta \gamma_0^3}{2} (p_T - v_0 m_T) + O(\delta v^2)$
 $f \sim \exp[-\beta E(v_0)].$

– Tsallis–Pareto: $\longrightarrow v_n^{\text{TS}} \approx \frac{\delta v_n \beta \gamma_0^3}{2} \frac{p_T - v_0 m_T}{1 + (q-1)\beta \gamma_0 (m_T - v_0 p_T)} + O(\delta v^2)$
 $f \sim [1 + (q-1)\beta E(v_0)]^{-1/(q-1)}$

Connecting spectra and v_2

- Spectra originating from hadronic sources

$$\left. \frac{dN}{2\pi p_T dp_T dy} \right|_{y=0} = \int_0^{2\pi} \frac{d\varphi}{2\pi} p^0 \left. \frac{dN}{d^3p} \right|_{y=0} = \sum_{m=0}^{\infty} \frac{a_m}{m!} \frac{\partial^m}{\partial v_0^m} f[E(v_0)] \approx f[E(v_0)] + O(\delta v^2)$$

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- Using the soft+hard model:

$$v_2 = \frac{w_{hard} f_{hard} + w_{soft} f_{soft}}{f_{hard} + f_{soft}} \quad \text{with the coefficient} \quad w_i = \frac{\delta v_i \gamma_i^3}{2T_i} \frac{p_T - v_i m_T}{1 + \frac{q_i - 1}{T_i} [\gamma_i(m_T - v_i p_T) - m]}$$

Connecting spectra and v_2

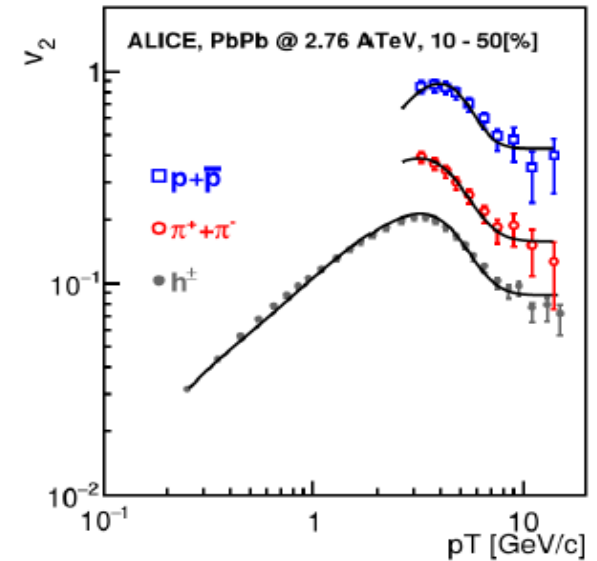
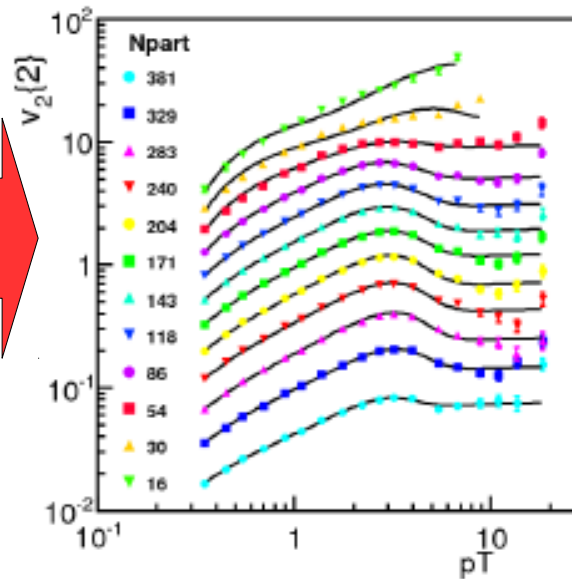
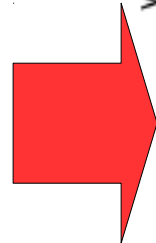
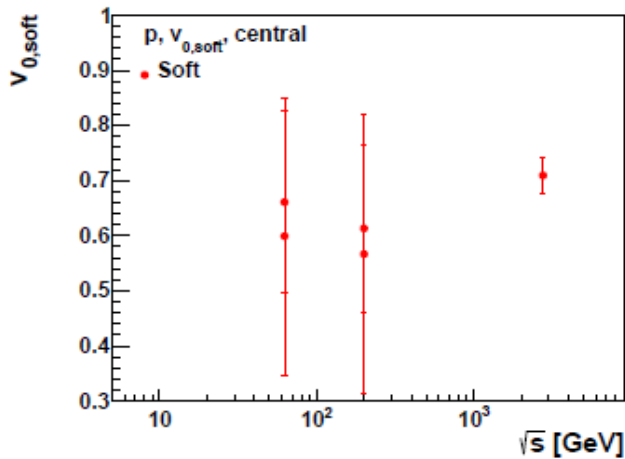
- Using the soft+hard model:

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with the coefficient

$$w_i = \frac{\delta v_i \gamma_i^3}{2T_i} \frac{p_T - v_i m_T}{1 + \frac{q_i - 1}{T_i} [\gamma_i(m_T - v_i p_T) - m]}$$

- Assuming v_0 only for the soft component v_2 can be obtained



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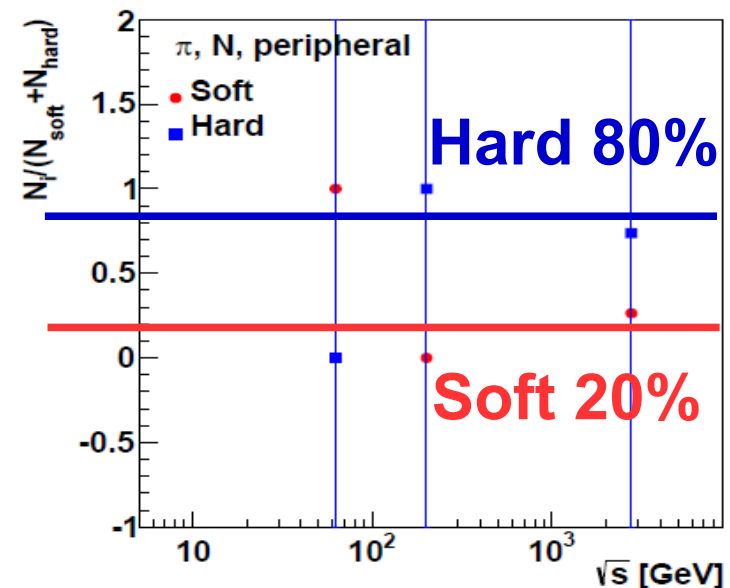
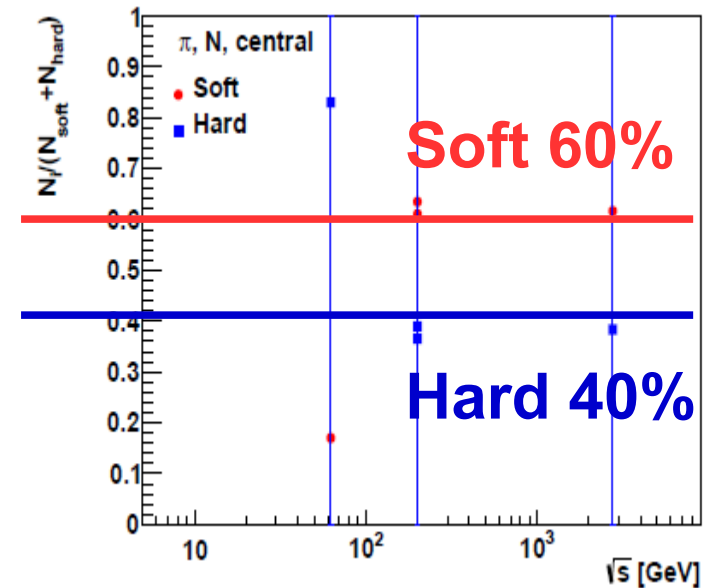
S U M M A R Y

- Non-extensive statistical approach in e^+e^- & pp
 - Obtained Tsallis/Rényi entropies from the first principles.
 - Providing physical meaning of $q=1-1/C + \Delta T^2/T^2$
 - *Boltzmann Gibbs limit $C \rightarrow \infty$ & $\Delta T^2/T^2 \rightarrow 0$ ($q \rightarrow 1$),*
 - *Tsallis – Pareto fits on spectra in e^+e^- , pp*
 - *Not working for larger system, like pA , AA and no flow.*
- Application of 'soft+hard' model in AA
 - Tsallis – Pareto + Exp does not working.
 - Double Tsallis – Pareto measures non-extensivity
 - **SOFT: $q \rightarrow 1$, suggest Boltzmann Gibbs (QGP)**
 - **HARD: $q > 1.1$, Tsallis – Pareto like**
 - Asimuthal anisotropy can be obtained too.

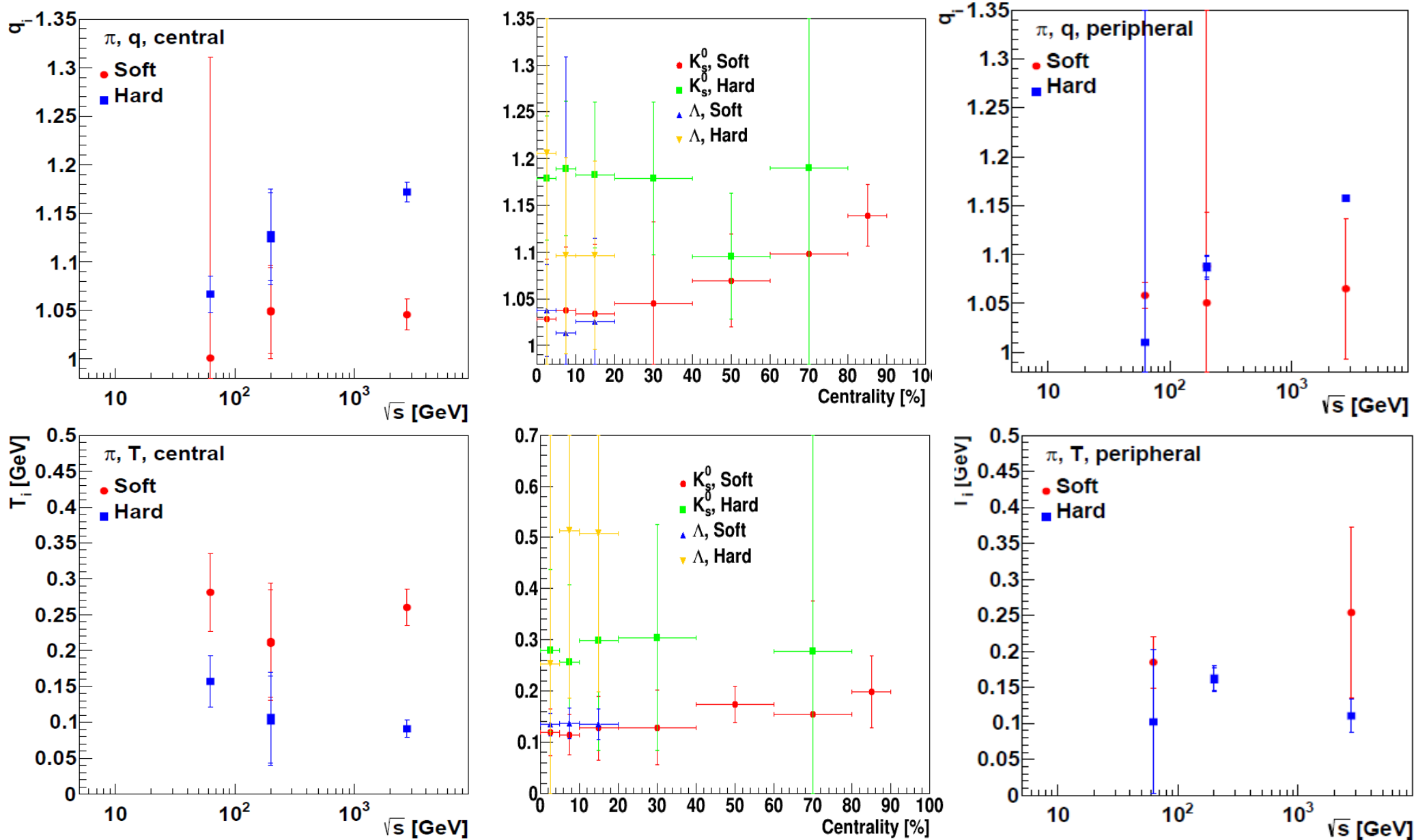
BACKUP

The c.m. Energy Dependence of N_{soft} & N_{hard}

- Energy dependence N_i/N_{tot}
 - Central
 - LHC: **HARD 40% + SOFT 60%**
 - RHIC: **HARD 80% + SOFT 20%**
 - Peripheral
 - LHC: **HARD 80% + SOFT 20%**
 - RHIC: **HARD 10% + SOFT 90%**



The c.m. Energy Dependence of q & T



Related publications..

1. arXiv:1409.5975: Statistical Power Law due to Reservoir Fluctuations and the Universal Thermostat Independence Principle
2. arXiv:1405.3963 Disentangling Soft and Hard Hadron Yields in PbPb Collisions at $\sqrt{s_{NN}} = 2.76$ ATeV
3. arXiv:1405.3813 New Entropy Formula with Fluctuating Reservoir, Physica A (in Print) 2014
4. arXiv:Statistical Power-Law Spectra due to Reservoir Fluctuations
5. arXiv:1209.5963 Nonadditive thermostatistics and thermodynamics, Journal of Physics, Conf. Ser. V394, 012002 (2012)
6. arXiv:1208.2533 Thermodynamic Derivation of the Tsallis and Rényi Entropy Formulas and the Temperature of Quark-Gluon Plasma, EPJ A 49: 110 (2013)
7. arXiv:1204.1508 Microcanonical Jet-fragmentation in proton-proton collisions at LHC Energy, Phys. Lett. B, 28942 (2012)
8. arXiv:1101.3522 Pion Production Via Resonance Decay in a Non-extensive Quark-Gluon Medium with Non-additive Energy Composition Rule
9. arXiv:1101.3023 Generalised Tsallis Statistics in Electron-Positron Collisions, Phys.Lett.B701:111-116,2011
10. arXiv:0802.0381 Pion and Kaon Spectra from Distributed Mass Quark Matter, J.Phys.G35:044012,2008

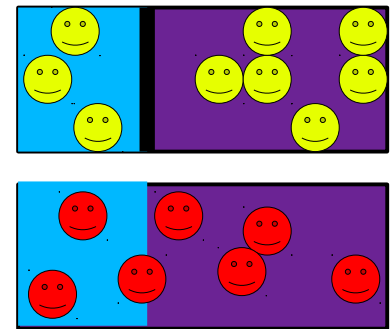
General derivation as improved canonical

The story is about...

- Two body thermodynamics:
1 subsystem (E_1) + one reservoir ($E-E_1$)
- Finite system, finite energy \rightarrow microcanonical description

- microcanonical $\sum_j \epsilon_j = E$

- canonical $\sum_j \langle \epsilon_j \rangle = E$



- Maximize a monotonic function of the Boltzmann-Gibbs entropy, $L(S)$ (0th law of thermodynamics)
- Taylor expansion of the $L(S) = \max$, principle beyond $-\beta E$

Description of a system & reservoir

- For generalized entropy function $L(S_{12}) = L(S_1) + L(S_2)$
- In order to exist β of the system $L(S(E_1)) + L(S(E - E_1)) = \max$

TS Biró P. Ván: Phys Rev. E84 19902 (2011)

- Thermal contact between system (E_1) & reservoir ($E - E_1$), requires to eliminate E_1 :

$$\begin{aligned}\beta_1 &= L'(S(E_1)) \cdot S'(E_1) \\ &= L'(S(E - E_1)) \cdot S'(E - E_1)\end{aligned}$$

- This is usually handled in canonical limit, but now, we keep **higher orders** in the Taylor-expansion in E_1/E

$$\beta_1 = L'(S(E)) \cdot S'(E) - [S'(E)^2 L''(S(E)) + S''(E) L'(S(E))] E_1 + \dots$$

Description of a system & reservoir

- Assuming $\beta_1 = \beta$, the Lagrange multiplier become familiar for us:

$$\beta = L'(S(E)) \cdot S'(E) = L'(S) \cdot \frac{1}{T}$$

- To satisfy this, need simply to solve

$$\frac{L''(S)}{L'(S)} = -\frac{S''(E)}{S'(E)^2}$$

- Universal Thermostat Independence (UTI)*
Principle: l.h.s. must be as an S-independent constant for solving $L(S)$,

$$\frac{L''(S)}{L'(S)} = a$$

- Based on $L(S) \rightarrow S$ for small S, coming from 3rd law of the thermodynamics
 $L'(0)=1$ and $L(0)=0$

$$L(S) = \frac{e^{aS} - 1}{a}$$

- EoS derivatives do have physical meaning:

$$\begin{aligned} S'(E) &= 1/T \\ S''(E) &= -1/CT^2 \end{aligned}$$

Description of a system & reservoir

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$$L(S) = \frac{e^{aS} - 1}{a}$$

- Simply the heat capacity of the reservoir:

$$a = 1/C$$

From two system to many...

- Analogue to Gibbs ensemble generalize

$$S = - \sum_i P_i \ln P_i \quad \rightarrow \quad L(S) = \sum_i P_i L(-\ln P_i)$$

- The L -additive form of a generally non-additive entropy, given by:

$$L(S(E_1)) - \beta E_1 = \frac{1}{a} \left(e^{aS(E_1)} - 1 \right) - \beta E_1 = \max.$$

- Introducing $a = 1/C(E)$ $\rightarrow L(S(E_1)) = L(-\ln P_1) = \frac{1}{a} (P_1^{-a} - 1)$

- we need to maximize:

$$\frac{1}{a} \sum_i (P_i^{1-a} - P_i) - \beta \sum_i P_i E_i - \alpha \sum_i P_i = \max.$$

which, results Tsallis:
and its inverse Rényi:

$$S_{\text{Tsallis}} := L(S) = \frac{1}{q-1} \sum_i (P_i - P_i^q)$$

$$S_{\text{Rényi}} := S = \frac{1}{1-q} \ln \sum_i P_i^q$$

The temperature slope

- Taking P_i weights of system, E_i , results cut power law:

$$P_i = \left(Z^{1-q} + (1-q) \frac{\beta}{q} E_i \right)^{\frac{1}{q-1}} = \frac{1}{Z} \left(1 + \frac{Z^{-1/C} e^{S/C} E_i}{C-1} \right)^{-C}$$

- Partition sum is related to Tsallis entropy, $L(S_1)$ and E_1

$$\ln_q Z := C \left(Z^{1/C} - 1 \right) = L(S_1) - \frac{1}{1-1/C} \beta E_1$$

- In $C \rightarrow \infty$ limit, the inverse log slope of the energy distribution:

$$T_{\text{slope}}(E_i) = \left(-\frac{d}{dE_i} \ln P_i \right)^{-1} = T_0 + E_i/C, \quad \text{with } T_0 = T e^{-S/C} Z^{1/C} (1 - 1/C)$$