

Hyper nucleus production in heavy ion collision experiments

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Strangeness in Quark Matter 2015
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Outline

Quick Introduction to
hypernuclear physics

Ultra relativistic region

relativistic region

Conclusion

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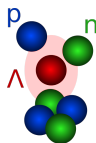
Conclusion

What and Why Hypernucleus ?

Classic example in Nuclear Physics:
Neutron Star

- ▶ Core of neutron star \rightarrow strangeness ?
- ▶ EoS of hyper-matter : Potential of hyperons
 - ▶ No direct study of hyperon-nucleon interaction

Yet hypernucleus can also be a probe to phase transition : QGP, Hadron gaz.



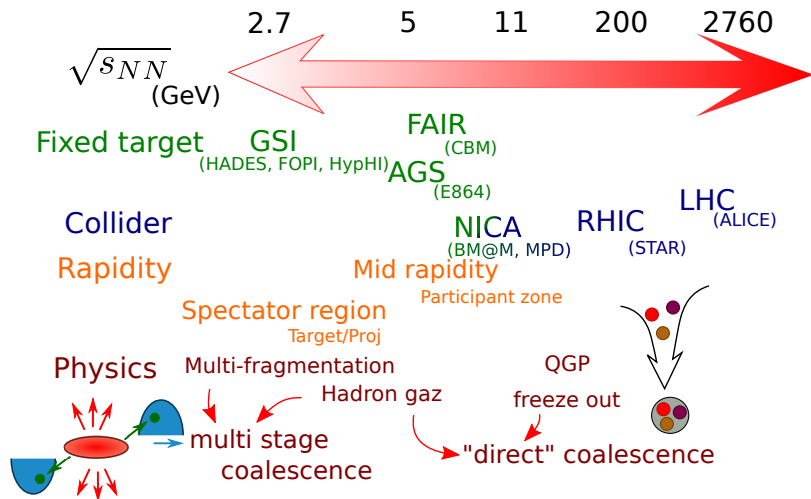
But, what is an hypernucleus ?

- ▶ Bound state of p,n and hyperon (Λ , Ξ , Σ) : A_Z

And How ?

- ▶ Common way : pion, kaon, e^- beam on fixed target.
- ▶ More recent : with heavy ion beam !

hypernuclei with ion beam



Observables on hypernuclear production in ion collisions

First "Classic" observables :

- ▶ Yields, Cross section, Multiplicity of the observed hypernuclei
- ▶ Absolute Or within Acceptance, also differential (y_0, Pt)
 - ▶ \rightarrow Useful to compare QMD models (dynamics of the reaction)
- ▶ Yield ratios : to Λ yield, to nuclear yield.
 - ▶ More sensitive to the dynamics than cross section.
 - ▶ Can be easier to extract from data (sometime eff. correction cancels out)

One special ratio : "Strangeness Population Factor" : Double ratio

$$S_3 = \frac{^3\Lambda\text{H}}{^3\text{He}} \cdot p/\Lambda$$

- ▶ Introduce by AGS (E864)
- ▶ With STAR, found out: nice probe to the phase transition and the QGP formation.
 - ▶ Theoretical calculation demonstration
[S. Zhang et al., Phys. Lett. B 684, 224 (2010)]

Outline

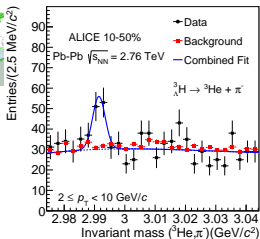
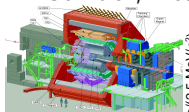
Quick Introduction to
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Ultra relativistic region

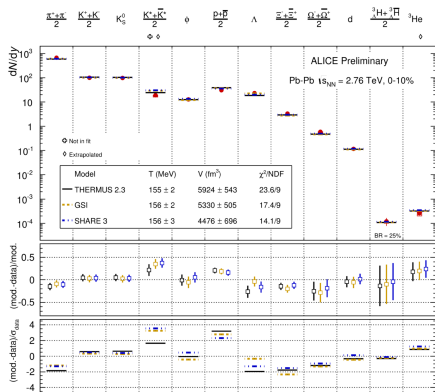
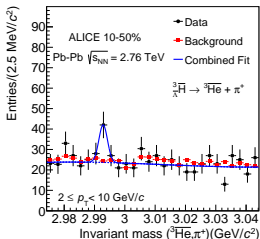
relativistic region

Conclusion

Mid-rapidity, freeze out : ALICE results

Collider, Reaction : Pb-Pb @ $\sqrt{s_{NN}} = 2.76$ TeV

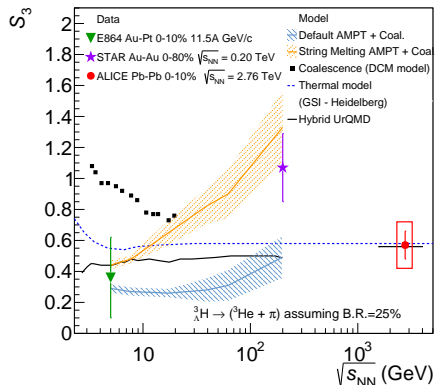
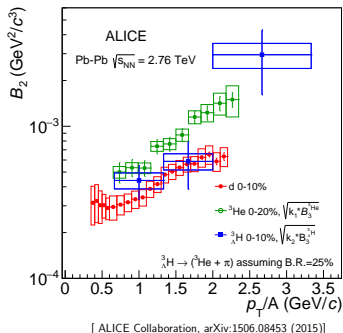
[ALICE Collaboration, arXiv:1506.08453 (2015)]



ALI-PREL-74463

[D. Elia, ALICE Collaboration, EPJ Conf. 95, 03008 (2015)]

Mid-rapidity, freeze out : ALICE results



$$B_A = E_A \frac{d^3 N_A}{d^3 p_A} / \left(E_P \frac{d^3 N_P}{d^3 p_P} \right)^A$$

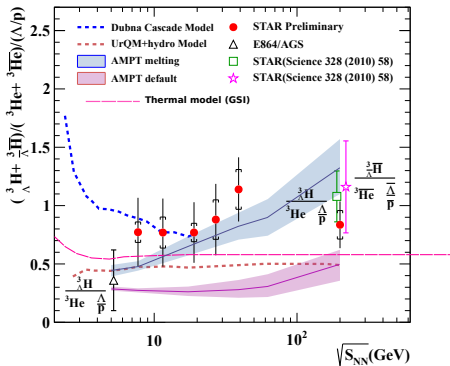
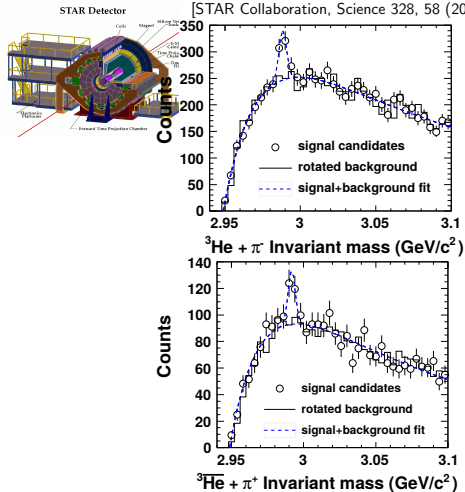
► Not as simple as expected !

AMPT + Coal. : S. Zhang et al., Phys.Lett.B684 (2010) 224.
 DCM model : J. Steinheimer et al., Phys.Lett.B714 (2012) 85.
 Thermal model : A. Andronic et al., Phys.Lett.B697 (2011) 203.
 Hybrid UrQMD : J. Steinheimer et al., Phys.Lett.B714 (2012) 85

Mid-rapidity, freeze out : STAR results

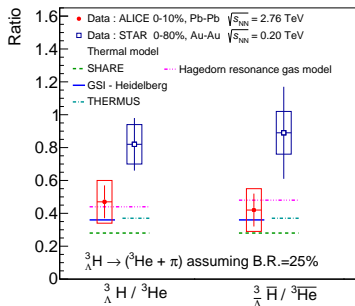
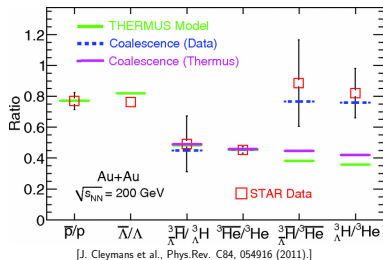
Collider, Reaction : Au+Au @ $\sqrt{s_{NN}} = 200$ GeV

[STAR Collaboration, Science 328, 58 (2010)]

[Y.G. Ma, STAR collaboration, EPJ Conf. 66, 04020 (2014)]
modified version !

Mid-rapidity, freeze out : STAR results

Coalescence or not coalescence ?



[ALICE collaboration, arXiv:1506.08453 (2015)]

- ▶ Coalescence: not clear anymore.
- ▶ Hadronization at the freeze-out.

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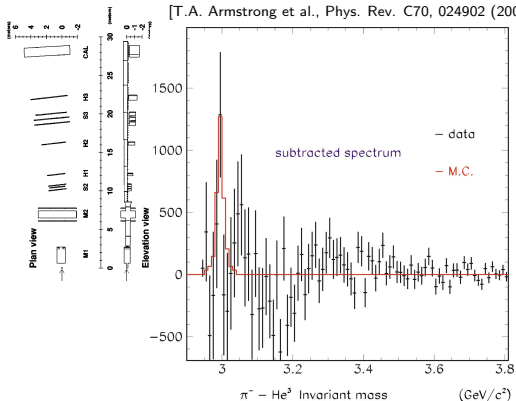
relativistic region

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Mid-rapidity, freeze out : AGS E864 results

Fixed target, Reaction: Au+Pt @ 11.5 AGeV/c or
 $\sqrt{s_{NN}} = 4.84 \text{ GeV}$

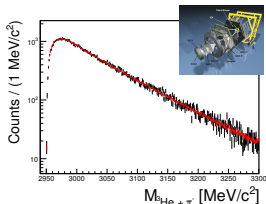
[T.A. Armstrong et al., Phys. Rev. C70, 024902 (2004)]



- ▶ ${}^3_{\Lambda}\text{H}$ signal : 2σ in rapidity range : [1.6, 2.6] ($y_{cm} \sim 2.3$)
- ▶ ${}^4_{\Lambda}\text{H}$: upper limit.
- ▶ First to use the ratio
 $S_3 = \frac{{}^3_{\Lambda}\text{H}}{{}^3\text{He}} \cdot p/\Lambda$ and
 $S_4 = \frac{{}^4_{\Lambda}\text{H}}{{}^4\text{He}} \cdot p/\Lambda$
- ▶ $S_3 = 0.36 \pm 0.26$
- ▶ $S_4(\text{upperLimit}) < 0.225$

Spectator region (target) : FOPI & HADES results :

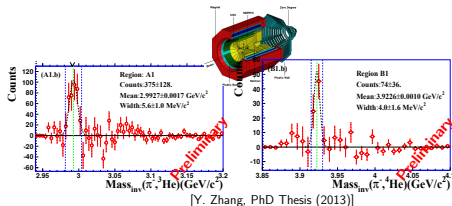
HADES: Ar+KCl @ 1.76 AGeV
or $\sqrt{s_{NN}} = 2.61$ GeV



[HADES Collaboration, Eur.Phys. J. A 49 (2013) 146]

- ▶ rapidity: $[-0.7, 0.1]$
($y_{cm} = 0.86$)
 - ▶ Upper limit :
 $\frac{^3\text{H}}{\Lambda} = 2.5 \pm 0.3 \cdot 10^{-2}$

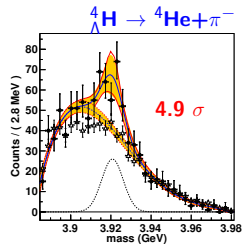
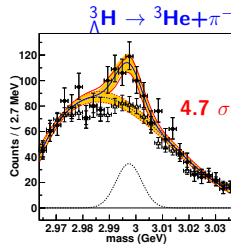
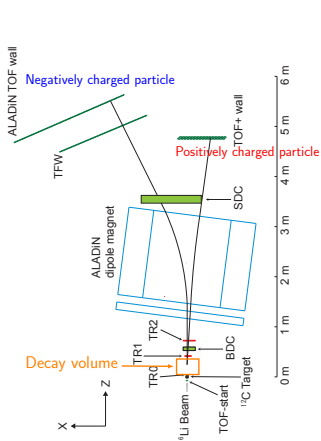
FOPI : Ni+Ni @ 1.91 AGeV or
 $\sqrt{s_{NN}} = 2.67$ GeV



- ▶ Rapidity : $[0.2, 0.4]$ ($y_{cm} = 0.9$)
& Pt/m $[0.2, 0.4]$
 - ▶ $\frac{^3\text{H}}{\Lambda} = 0.52 \pm 0.04$
- ▶ Rapidity : $[0.15, 0.4]$ ($y_{cm} = 0.9$)
& Pt/m $[0.15, 0.35]$
 - ▶ $\frac{^4\text{H}}{\Lambda} \sim 0.09$

Spectator region (projectile) : HypHI results

Fixed target, Reaction : ${}^6\text{Li} + {}^{12}\text{C} @ 2 \text{ AGeV}$ or $\sqrt{s_{NN}} = 2.7 \text{ GeV}$



[C. Rappold et al., Nucl. Phys. A. **913**, 170 (2013)]

► Before in proj. region :

- At LBNL Bevatron
 ${}^{16}\text{O} + \text{polyethylene} @ 2.1 \text{ AGeV}$

[K. Nield et al., Phys.Rev. C13, 1263 (1976)]

- At JINR Dubna, ${}^4\text{He} @ 3.7 \text{ AGeV}$ &
 ${}^7\text{Li} @ 3.0 \text{ AGeV} + \text{polyethylene}$

[A. U. Abdurakhimov et al., N.Cimento A102, 645 (1989)]

Spectator region (projectile) : HypHI results

Production cross section & multiplicity :
CS

$$\Lambda \quad 1.7 \pm 0.8 \text{ mb}$$

$${}^3_{\Lambda}\text{H} \quad 3.9 \pm 1.4 \text{ } \mu\text{b}$$

$${}^4_{\Lambda}\text{H} \quad 3.1 \pm 1.0 \text{ } \mu\text{b}$$

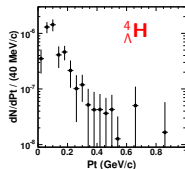
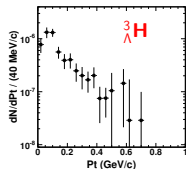
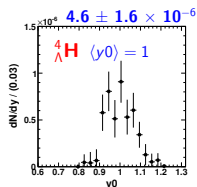
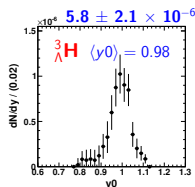
$${}^3_{\Lambda}\text{H}/{}^4_{\Lambda}\text{H} \quad 1.4 \pm 0.8$$

$${}^3_{\Lambda}\text{H}/\Lambda \quad 2.6 \pm 1.4 \times 10^{-3}$$

$${}^4_{\Lambda}\text{H}/\Lambda \quad 2.1 \pm 1.1 \times 10^{-3}$$

$${}^3_{\Lambda}\text{H}/{}^3\text{He} \cdot p/\Lambda \quad 0.28 \pm 0.14$$

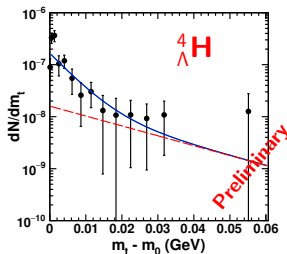
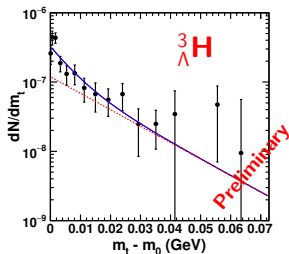
$${}^4_{\Lambda}\text{H}/{}^4\text{He} \cdot p/\Lambda \quad 0.08 \pm 0.04$$

[C. Rappold *et al.*, Phys.Lett. B747 (2015) 129]

Spectator region : HypHI results

Inverse slope T , m_t spectrum :

$$f(m_t - m_0) = K_1/T_1 e^{-(m_t - m_0)/T_1} + K_2/T_2 e^{-(m_t - m_0)/T_2}$$



- ▶ for ${}^3_{\Lambda}\text{H}$: $T_1 \sim 7 \pm 2$ MeV & $T_2 \sim 18 \pm 7$ MeV
- ▶ for ${}^4_{\Lambda}\text{H}$: $T_1 \sim 6 \pm 2$ MeV & $T_2 \sim 13 \pm 6$ MeV
- ▶ very similar to multi-fragmentation ALADIN results

[T. Odeh et al., Phys. Rev. Lett. 84 (2000) 4557]

- ▶ and Goldhaber model : [A.S. Goldhaber, Phys. Lett. B 53 (1974) 306]

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In ultra relativistic regime :

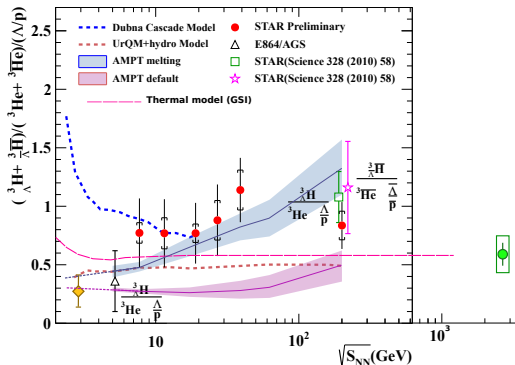
- ▶ From ALICE and STAR : similar tendency.
- ▶ yet sensitive observables give different conclusion on processes.
- ▶ Hadronization at the freeze-out vs coalescence
- ▶ Coalescence dependent to the volume and locality within QGP
- ▶ More precise measurements will be needed.

In relativistic regime :

- ▶ mid-rapidity: similar story: freeze-out.
- ▶ spectators: different production mechanism
 - ▶ coalescence of mid-rapidity Λ with spectator fragments.
 - ▶ FOPI : \sim target region / HypHI : spectator region
 - ▶ probe to what happen between the participants & spectator !
 - ▶ ex: Yield ratio sensitive to time interval of the interaction.

Conclusion

Mock-up figure of S_3 factor with all available results:



[T.A. Armstrong et al., Phys. Rev. C70, 024902 (2004)]

[C. Rappold et al., Phys.Lett. B747, 129 (2015)]

[Y.G. Ma, STAR collaboration, EPJ Conf. 66, 04020 (2014)]

[STAR Collaboration, Science 328, 58 (2010)]

[ALICE collaboration, arXiv:1506.08453 (2015)]

Conclusion

What next ?

- ▶ from ALICE & STAR :
 - ▶ Next hypernuclei, $A=4$ will be difficult to measure (penalty factor)
 - ▶ Might focus $A=3$ hypernuclei with different baryons.
- ▶ FOPI : work in progress for finalizing analysis.
- ▶ HypHI :
 - ▶ Second exp. performed $^{20}\text{Ne}+^{12}\text{C}$ @ 2 AGeV
 - ▶ spectroscopy: Λ analysis done ! $^3_{\Lambda}\text{H}$ and $^4_{\Lambda}\text{H}$ done !
 - ▶ $^7_{\Lambda}\text{Li}$ and $^6_{\Lambda}\text{He}$ analysis in progress.
 - ▶ reaction aspect: to be started.
 - ▶ Next exp. on proton rich hypernuclei.
- ▶ FAIR & NICA : excellent opportunities for the hypernuclear study !